



# THE NEW MOON

Jeff Plescia

The Johns Hopkins University

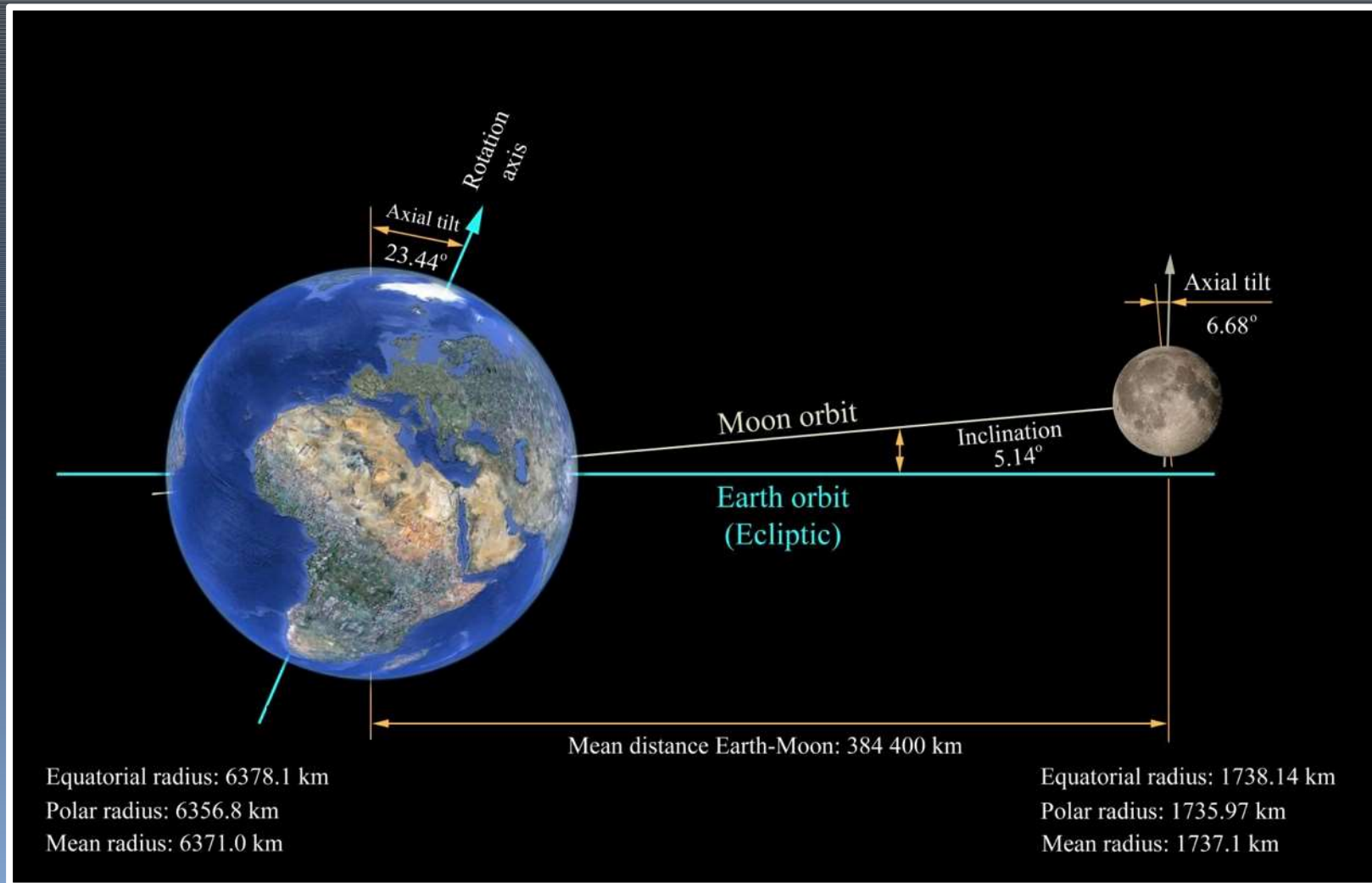
Applied Physics Laboratory

April 12, 2016

# The View From The Earth

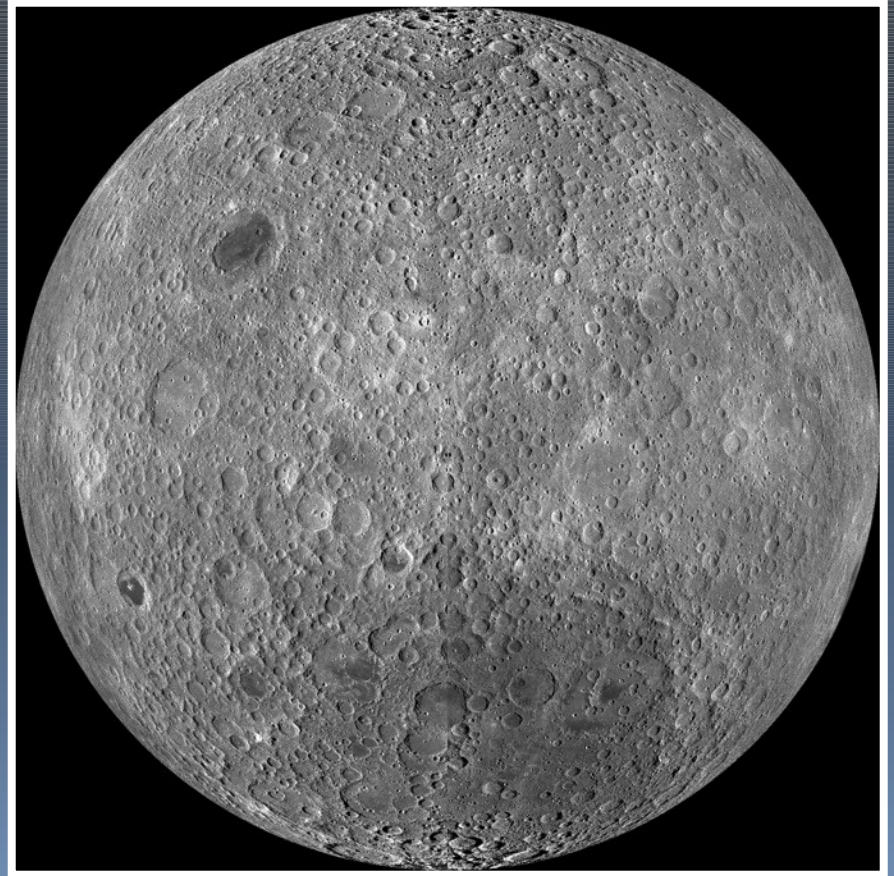
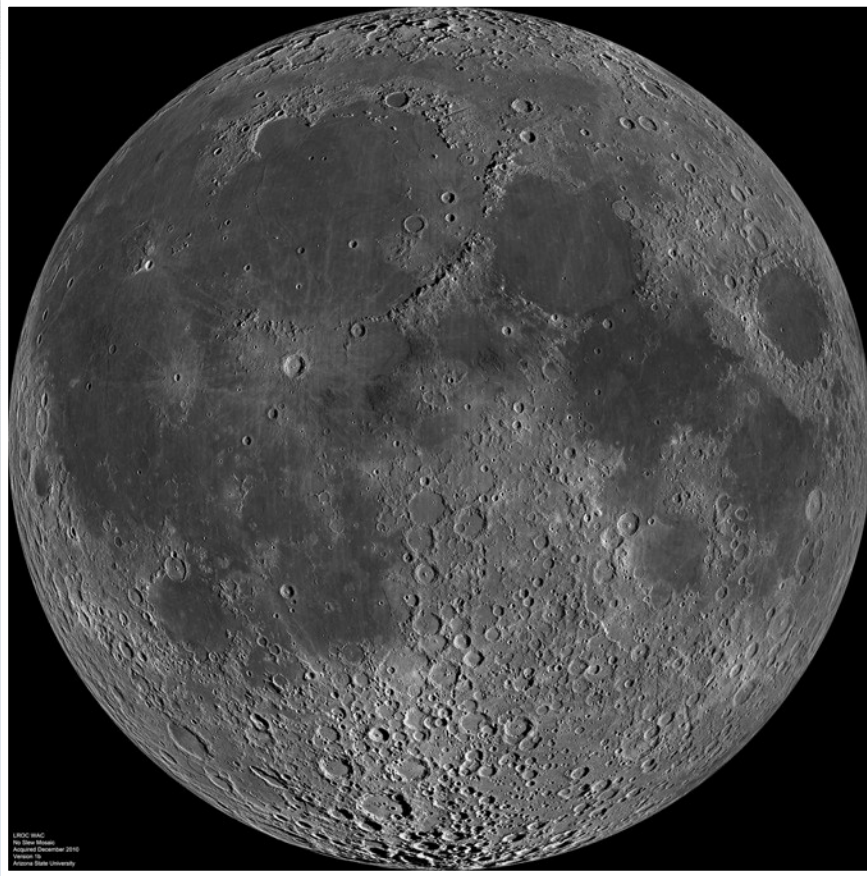


# Moon's Orbit





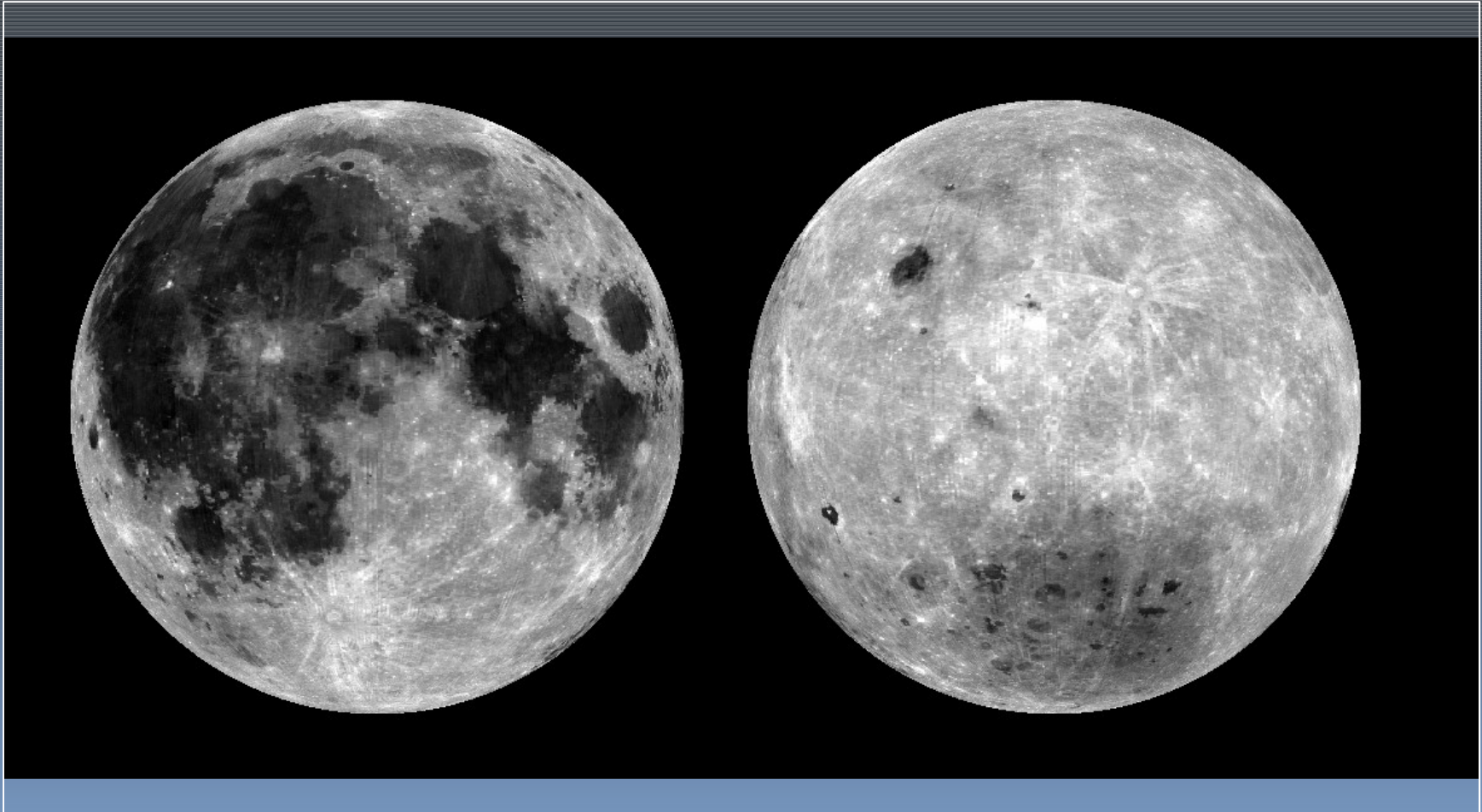
## Nearside – Farside - Sunlight



Every part of the Moon's surface was imaged at the same local time on the surface.  
You would never see this view naturally.

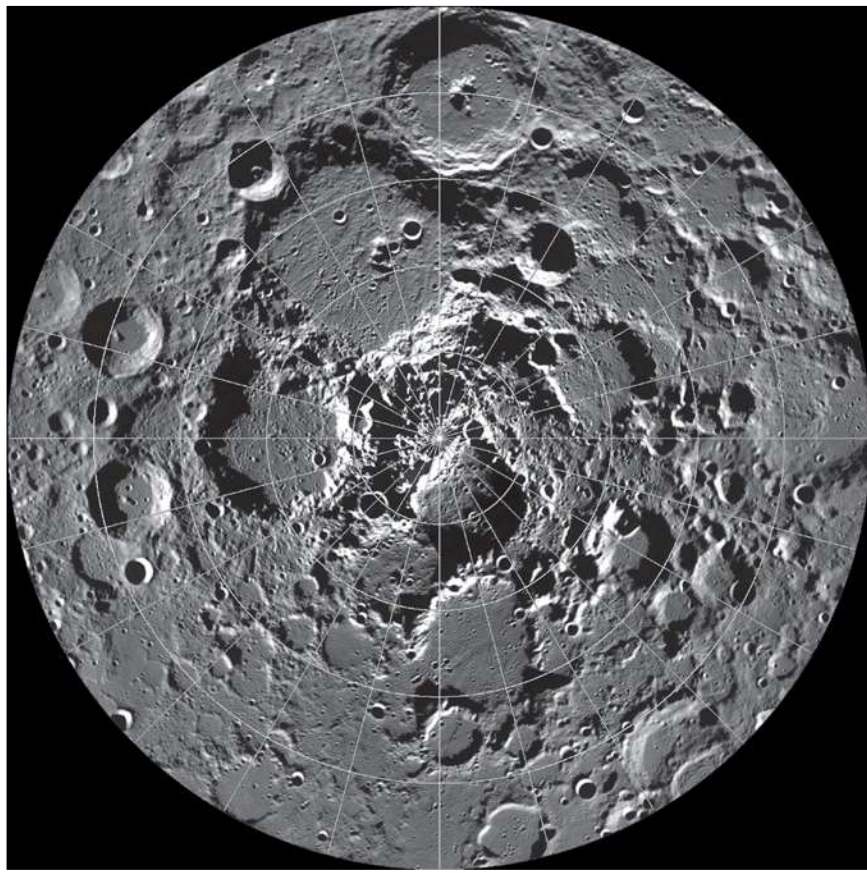


## Nearside – Farside - Laser

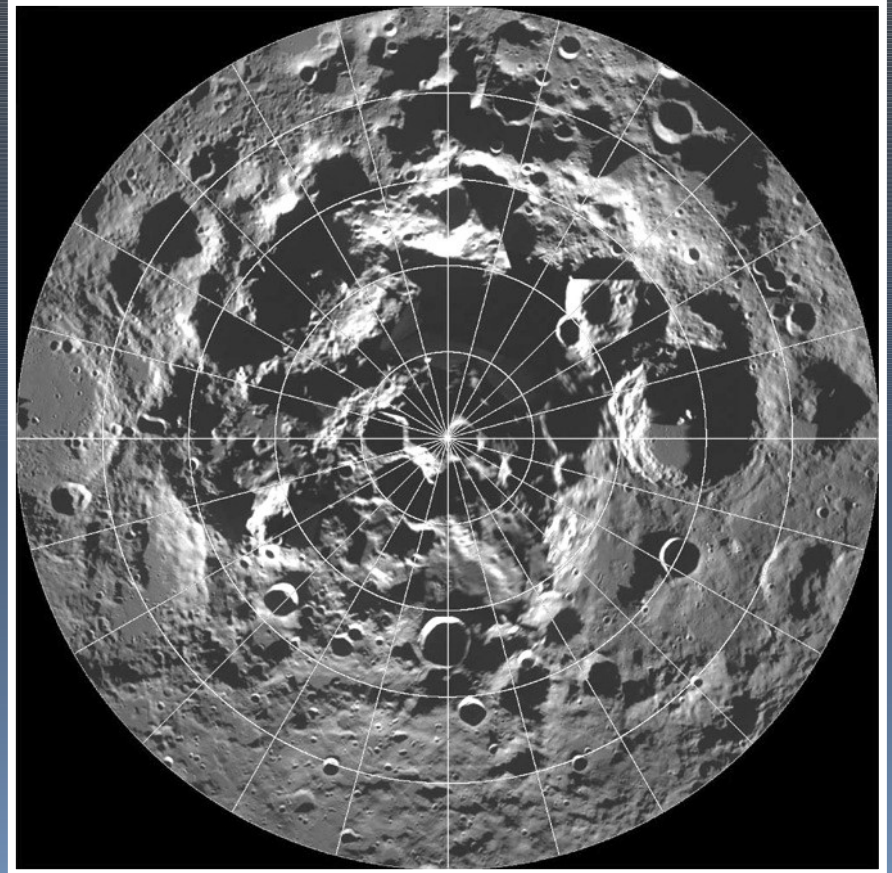


The laser beam, used to measure distance, reflects light back. This shows the surface illuminated by the laser from directly above the surface. Looks like the previous images.

## The Poles - Sunlight



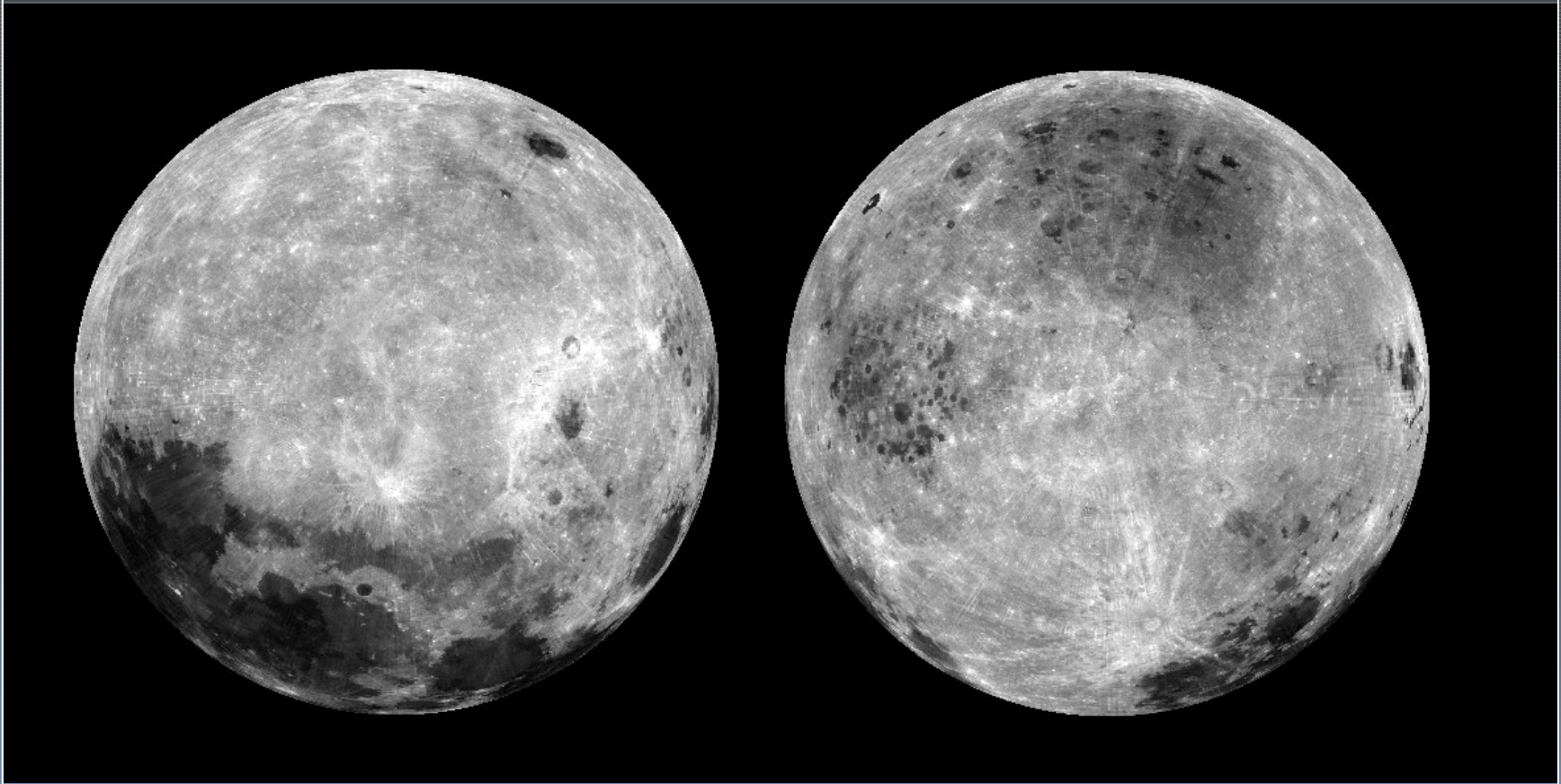
North Pole



South Pole

Every part of the Moon's surface was imaged at the same local time on the surface. You would never see this view naturally. But the Sun is always on the horizon – it never rises higher than  $1.5^\circ$  above the horizon. Because the Sun never rises very high, there are areas of permanent shadow.

## The Poles - Laser



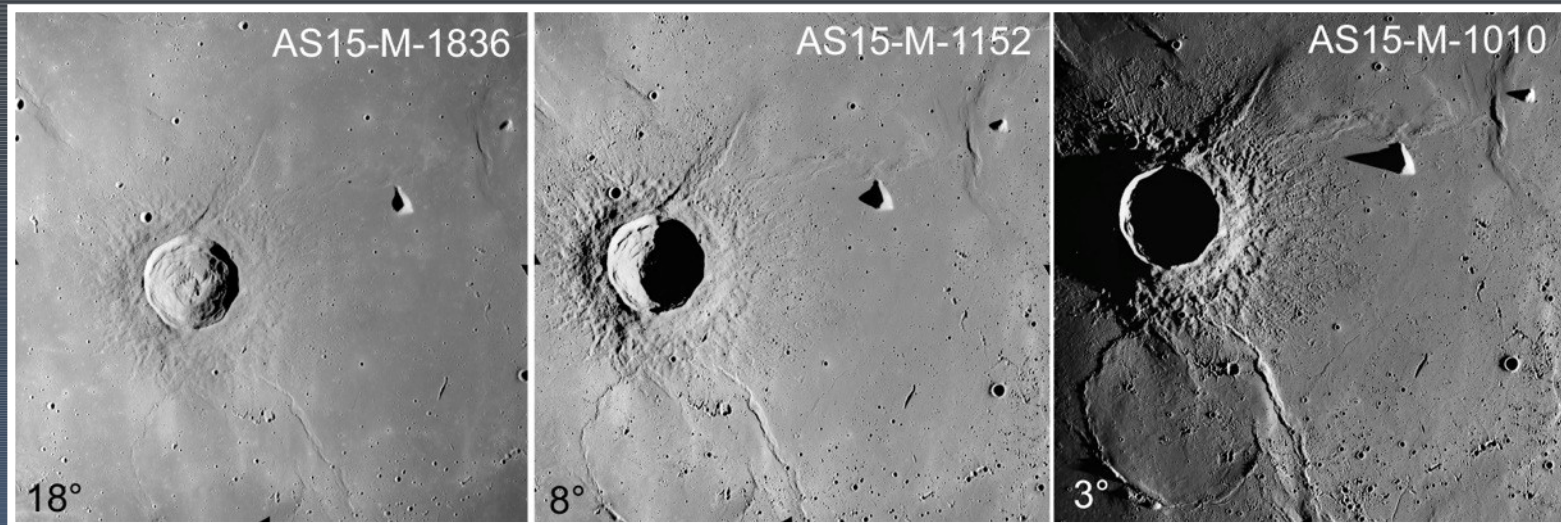
North Pole

South Pole

The laser beam, used to measure distance, reflects light back. This shows the surface illuminated by the laser from directly above the surface. Because the laser on the spacecraft, it illuminates the surface from above and the shadows disappear allowing us to see into the areas of permanent darkness.



# Illumination Effects

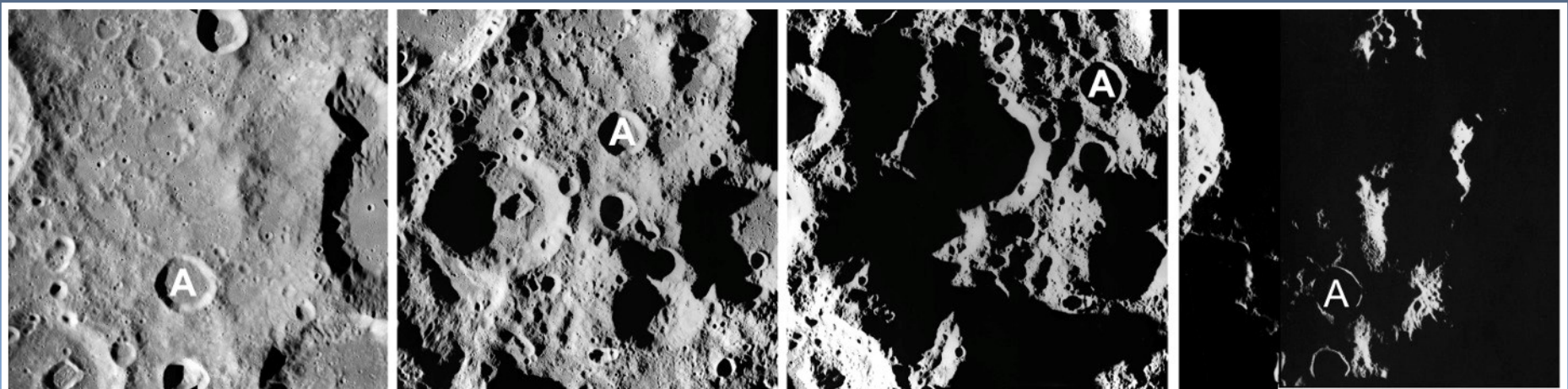


19°

8°

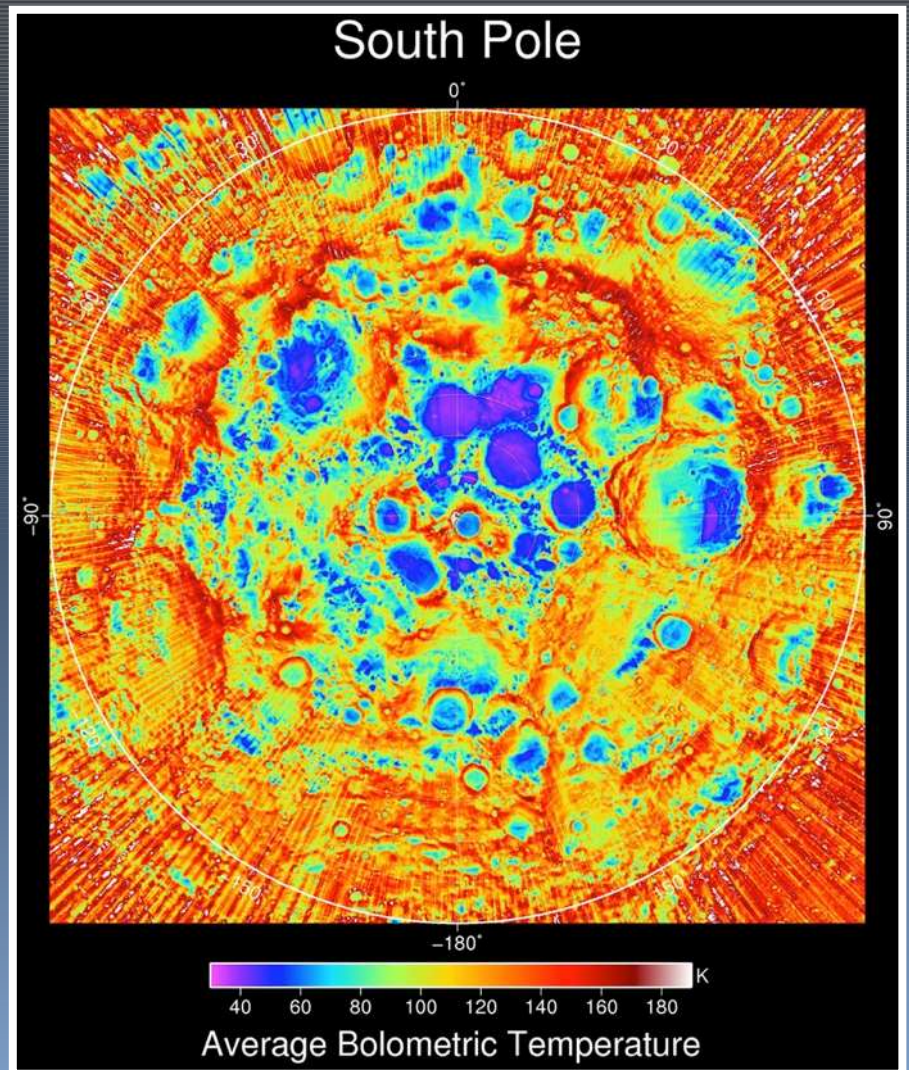
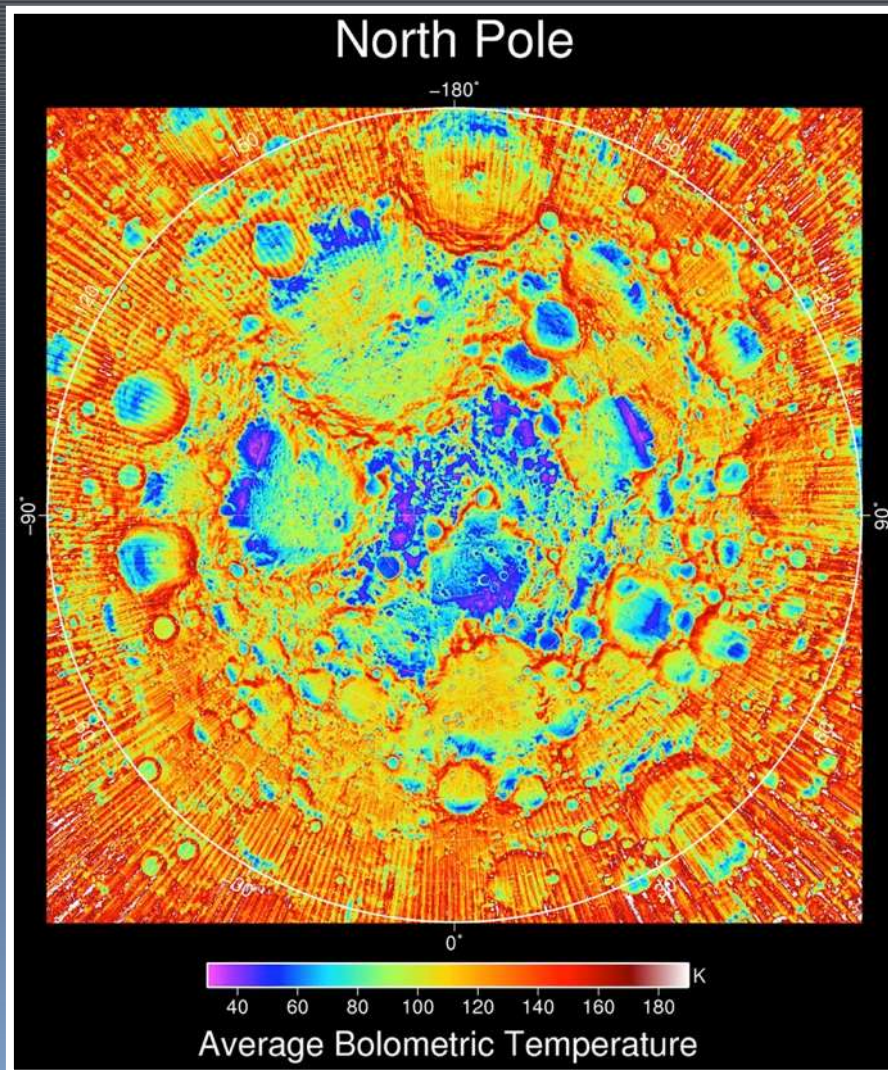
4°

1°





# Average Polar Temperature



*Polar Craters: 30K (-405° F) Pluto: 44 Kelvin (-380° F).*

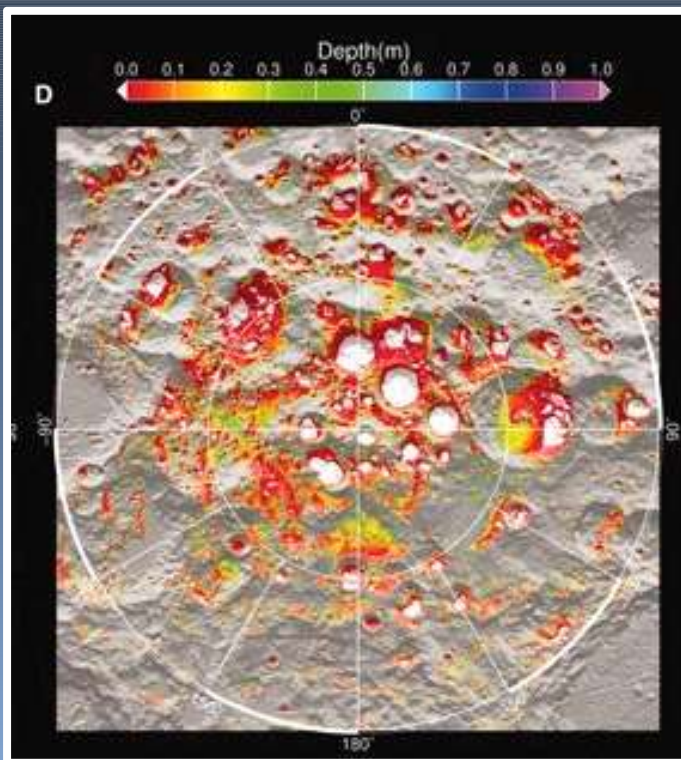


# Polar Temperatures

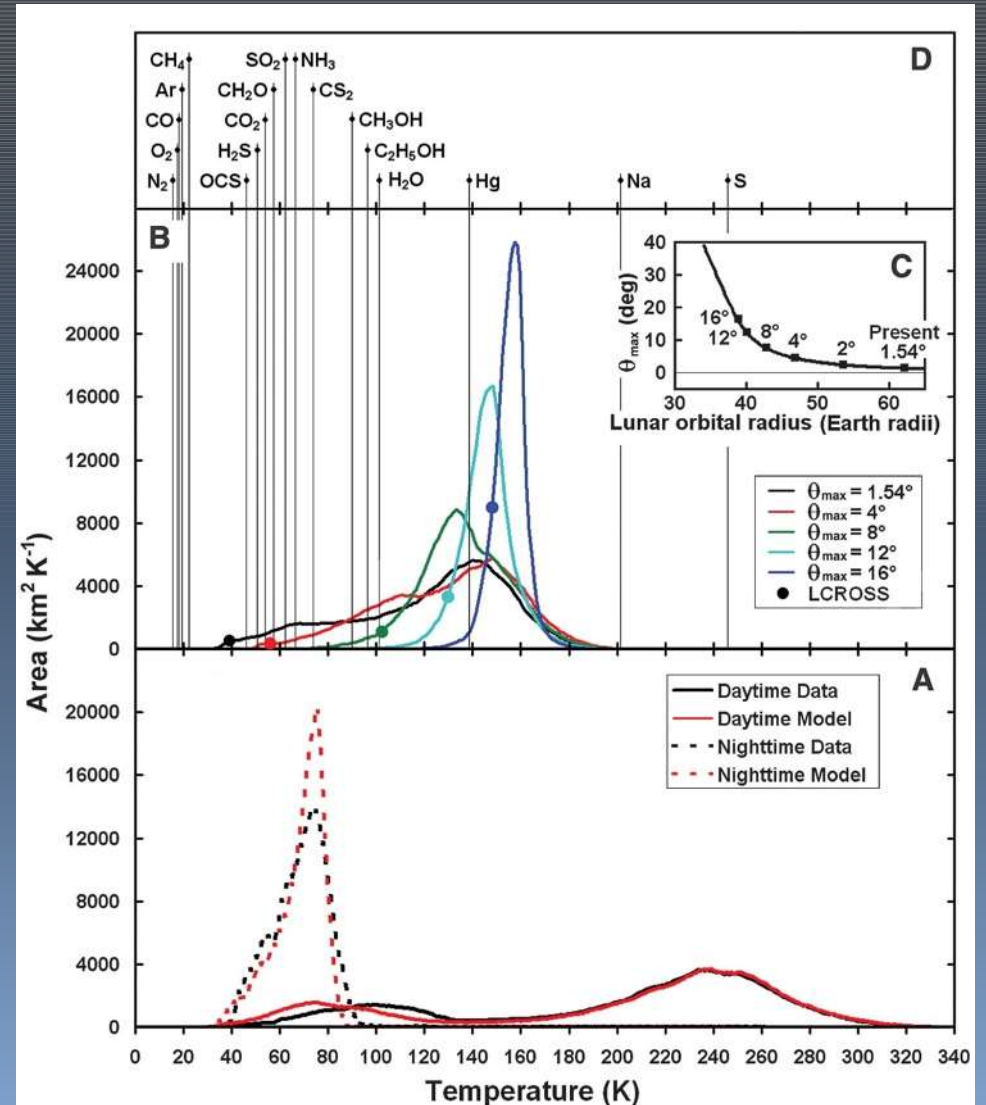
LRO Diviner temperature measurements and thermal analysis

Provides map of where different compounds are stable against sublimation.

Modeling predicts the depth to a thermally stable layer.

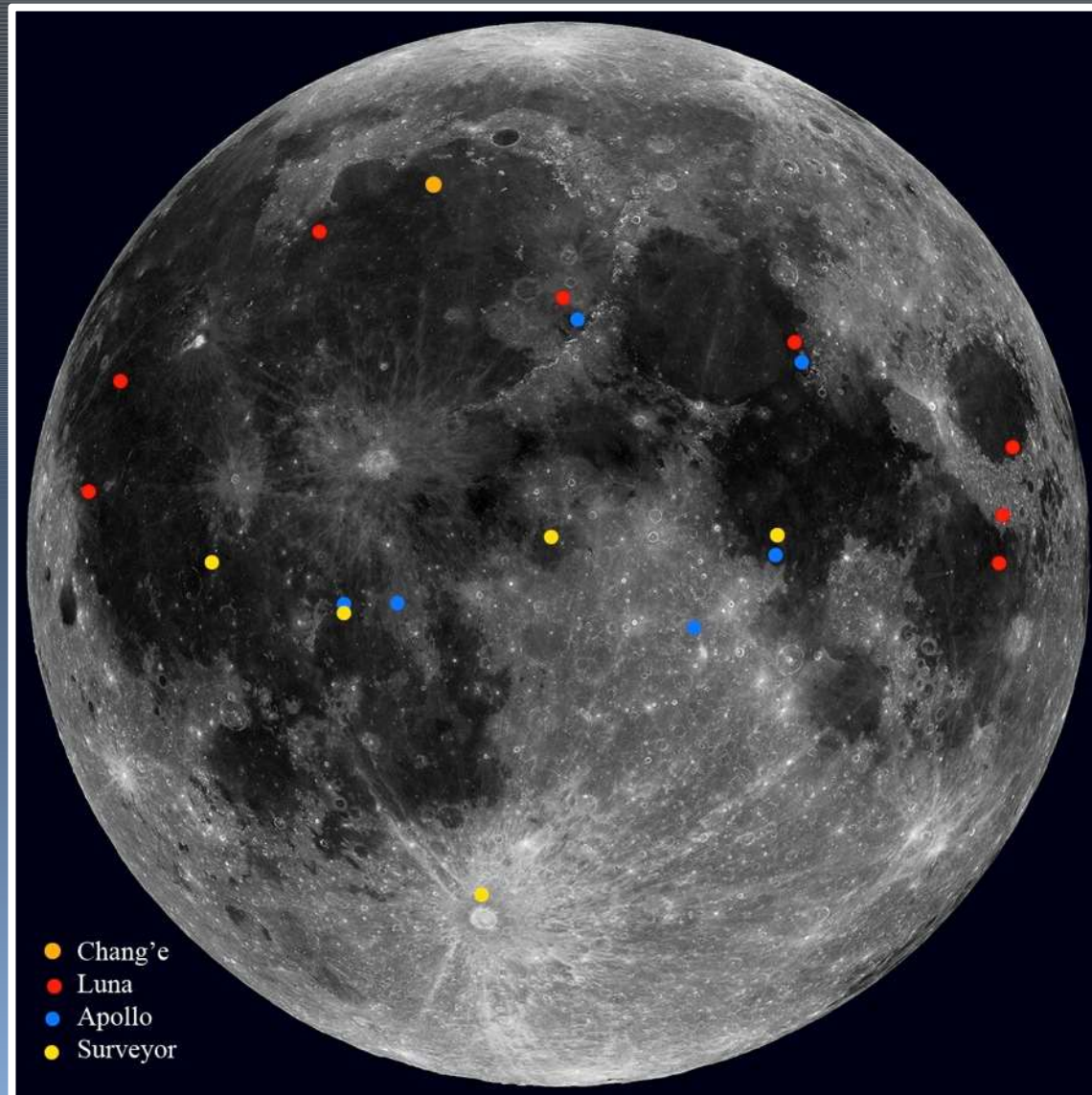


Depth to freezing

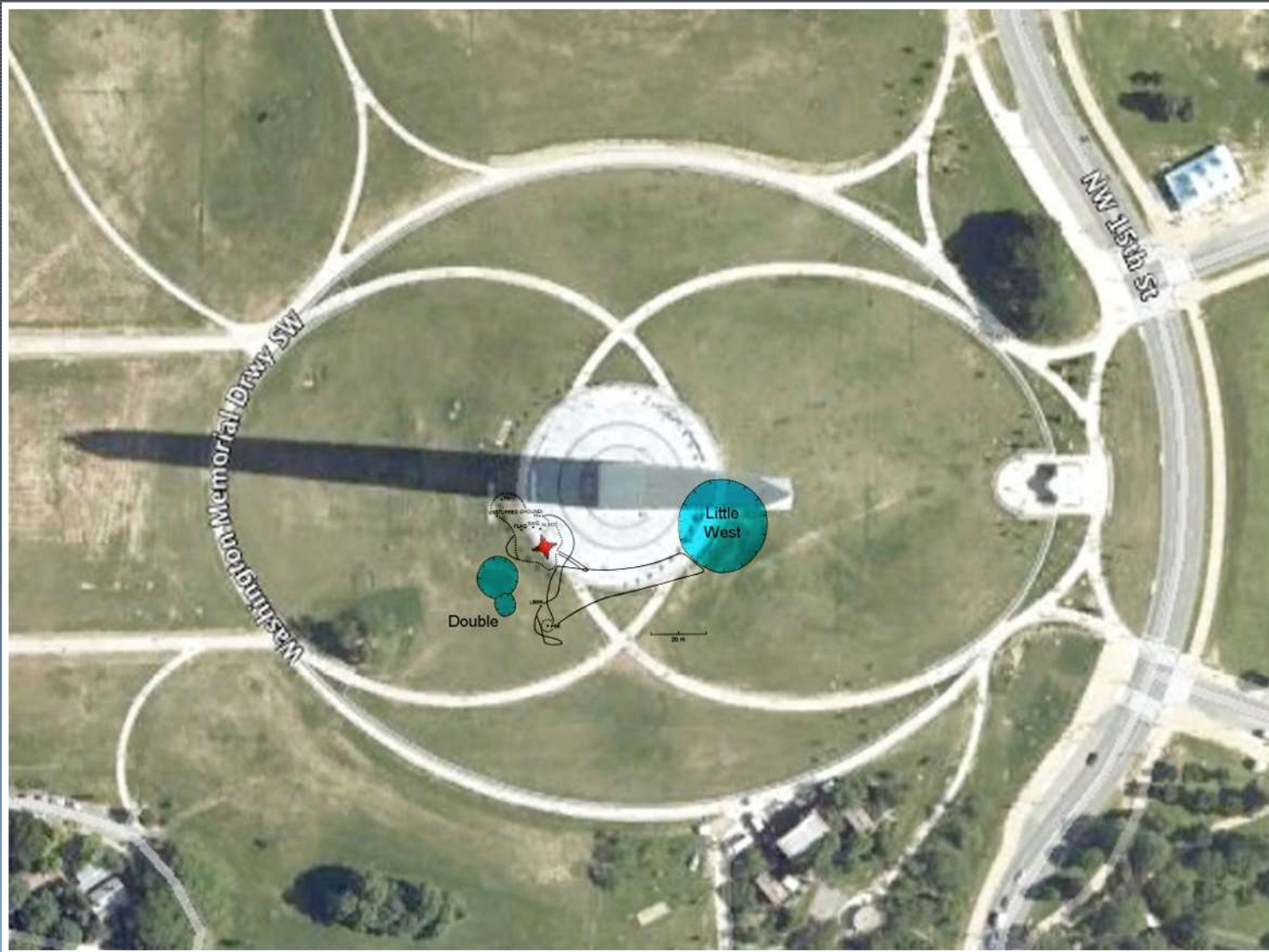




# Landing Sites

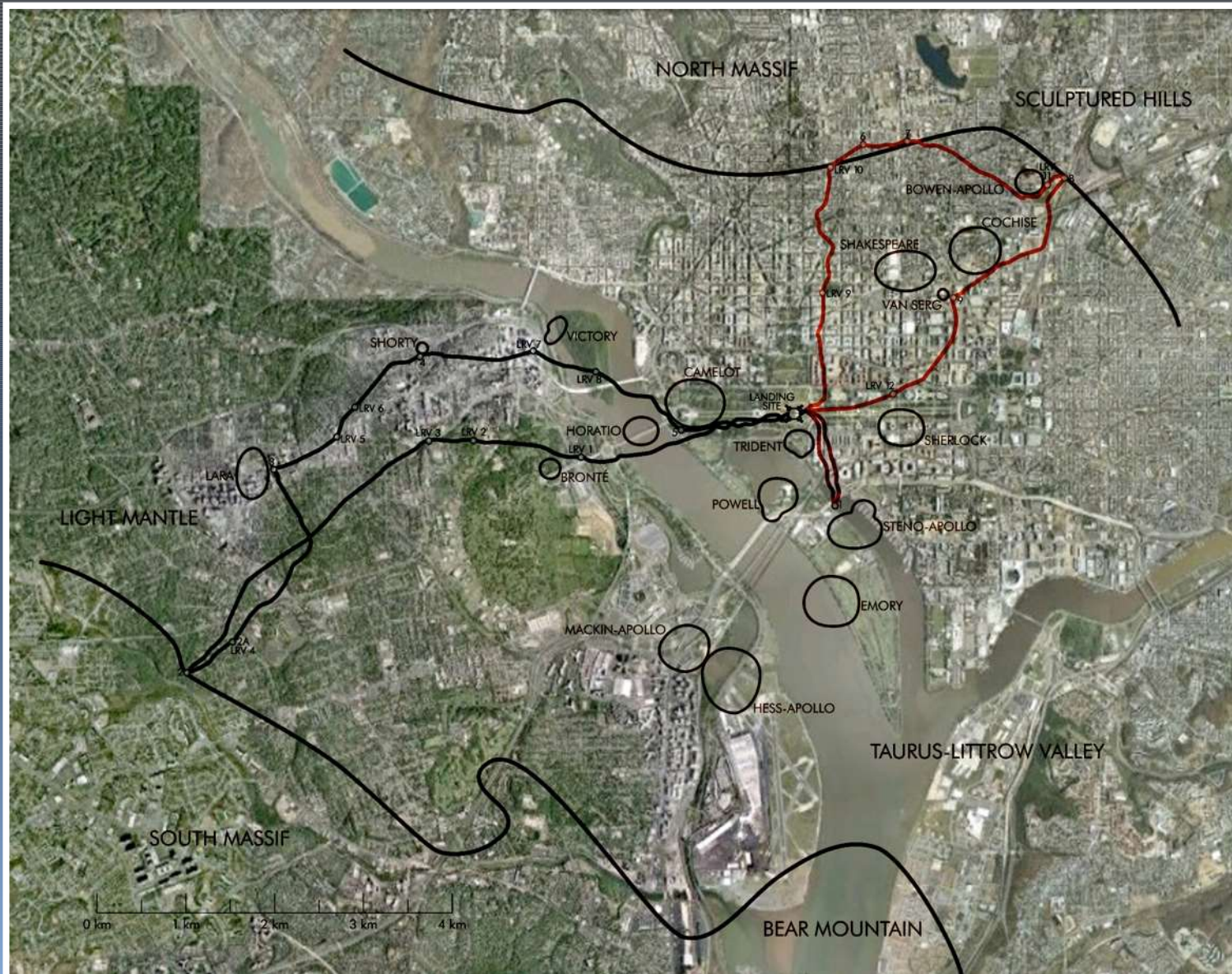


# Apollo 11 Traverse



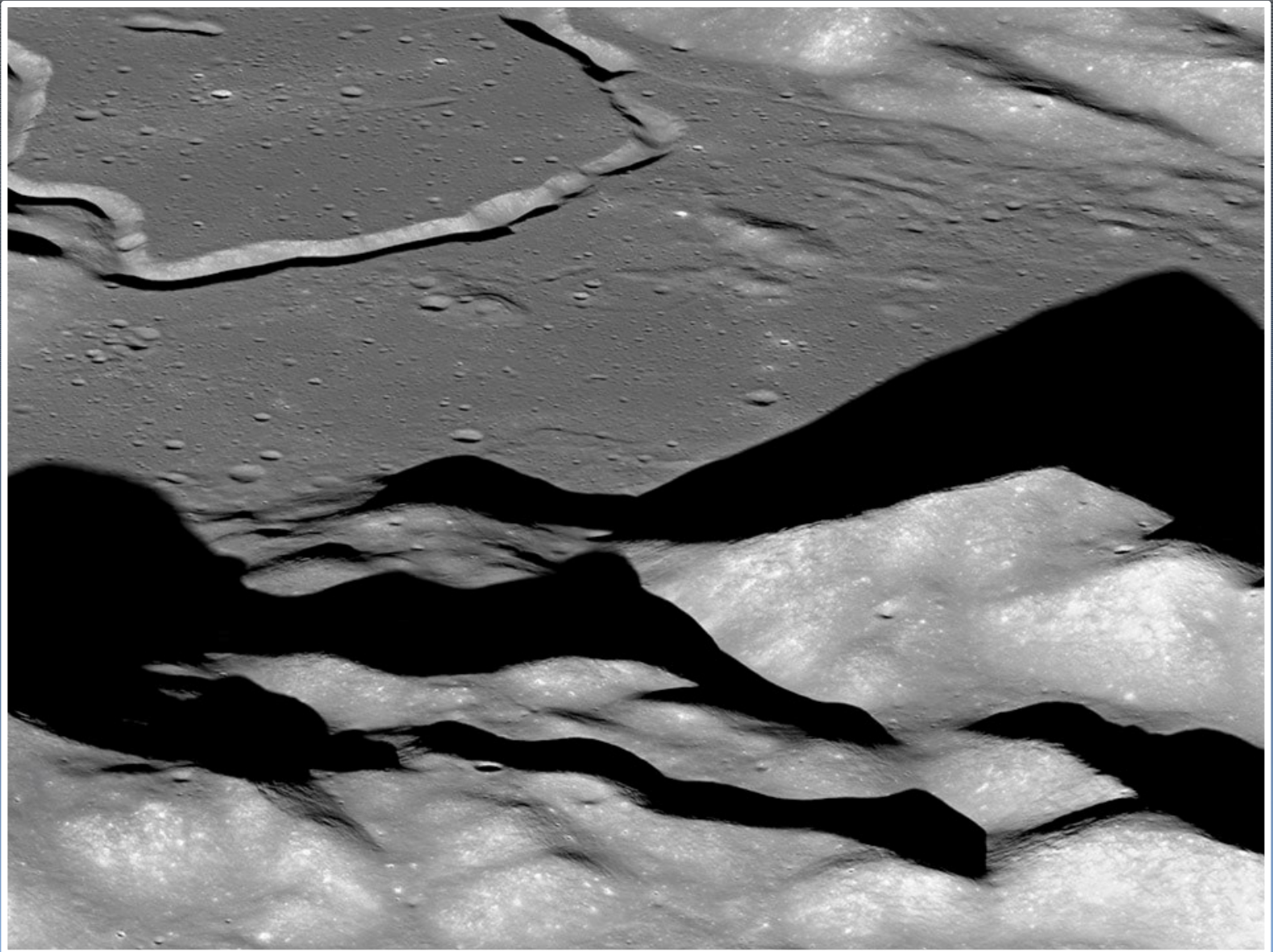


# Apollo 17 Traverse

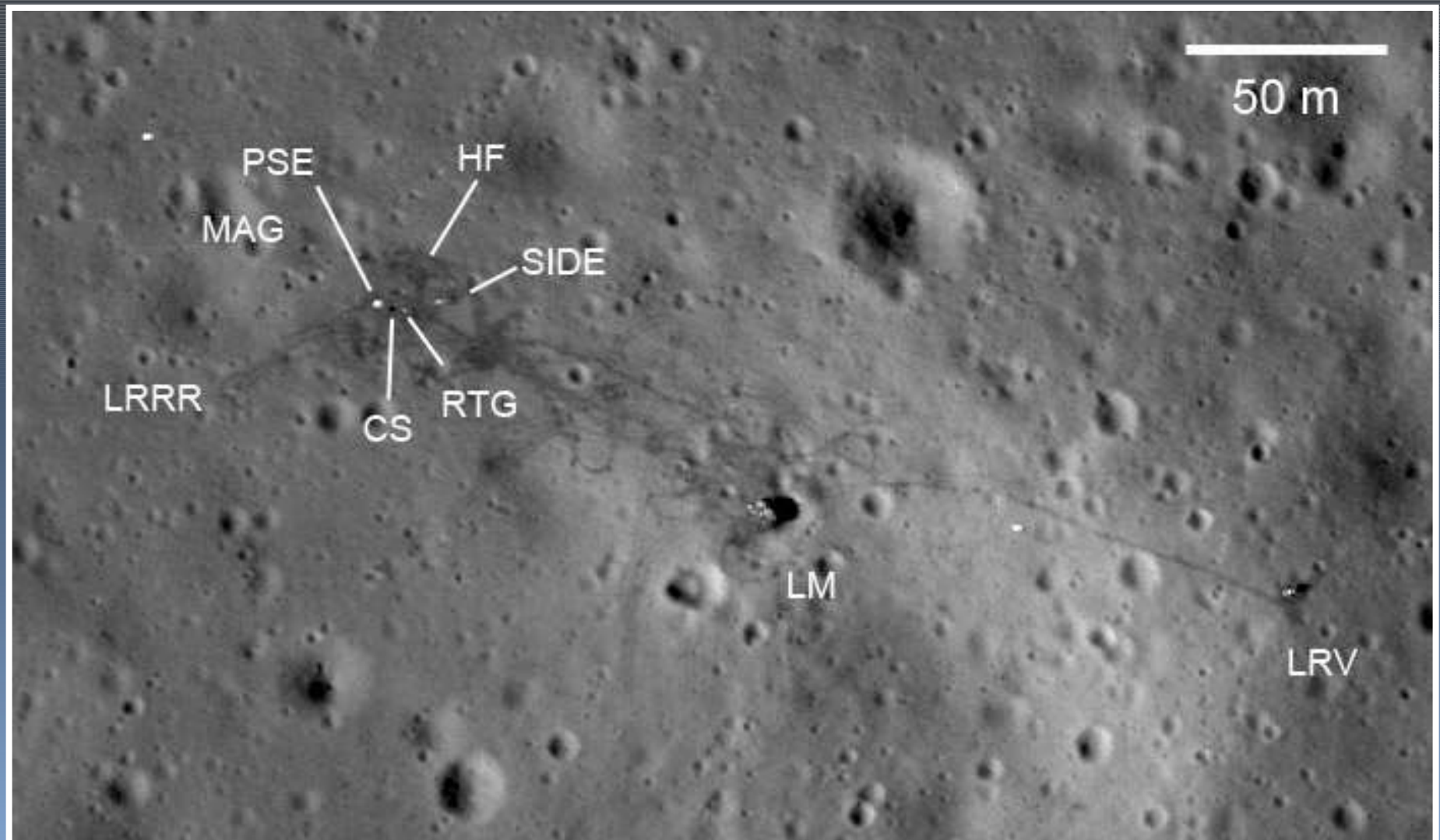




# Apollo 15



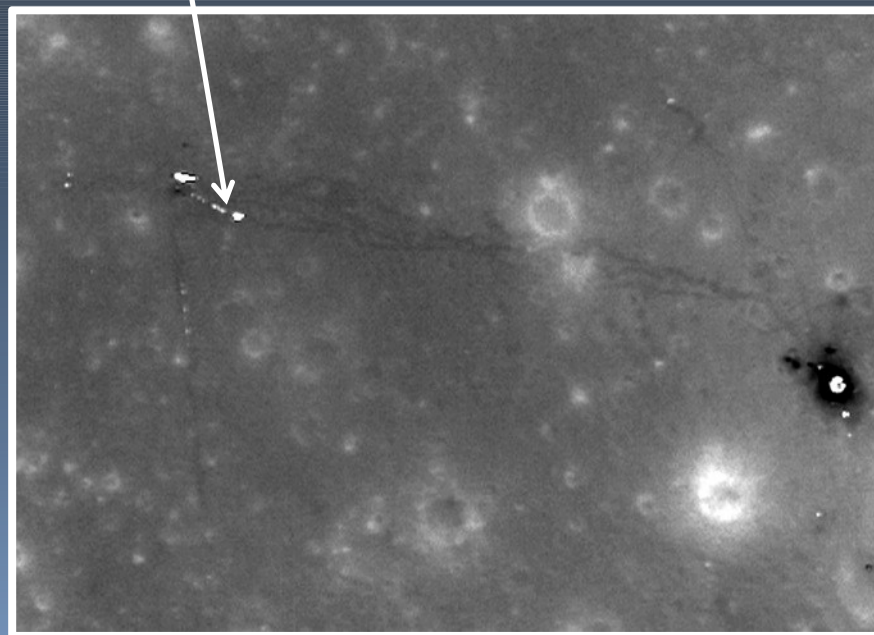
# Apollo 15



# Apollo 14 Site



Wire ribbon



The ribbon wire was laid out across the surface to connect experiments. It's flat shiny surface reflects sunlight, and is visible 45 year later.



# Apollo 17 – Station 6



# Saturn SIVB

Saturn IB (Stage 2)

Saturn V (Stage 3)

Height: 17.8 m

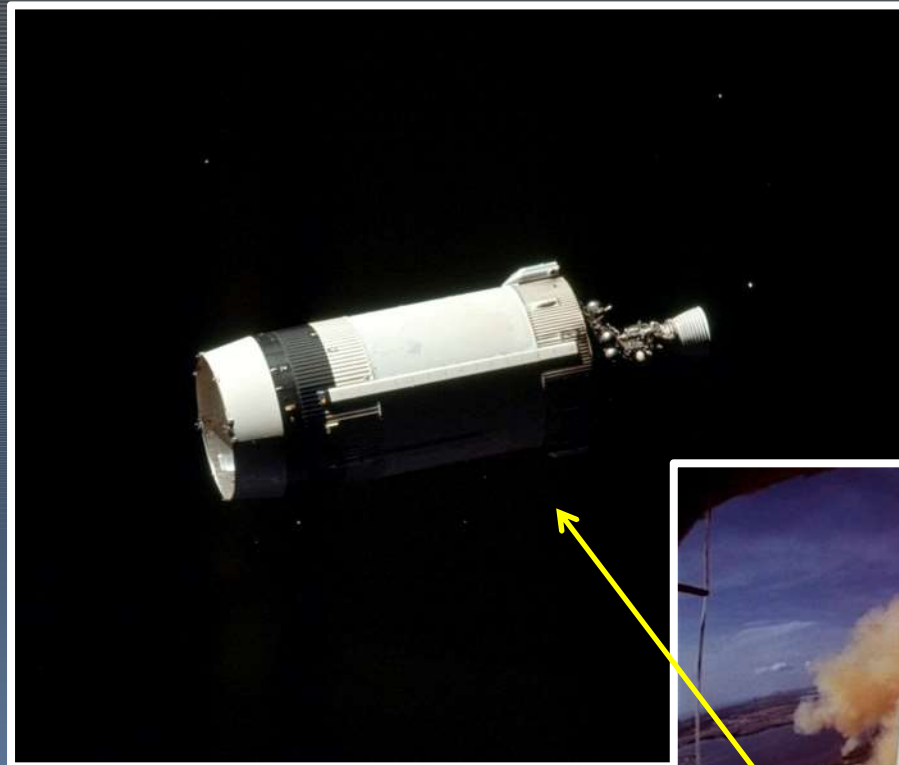
Diameter: 6.6 m

Mass (Wet): 119,900 kg

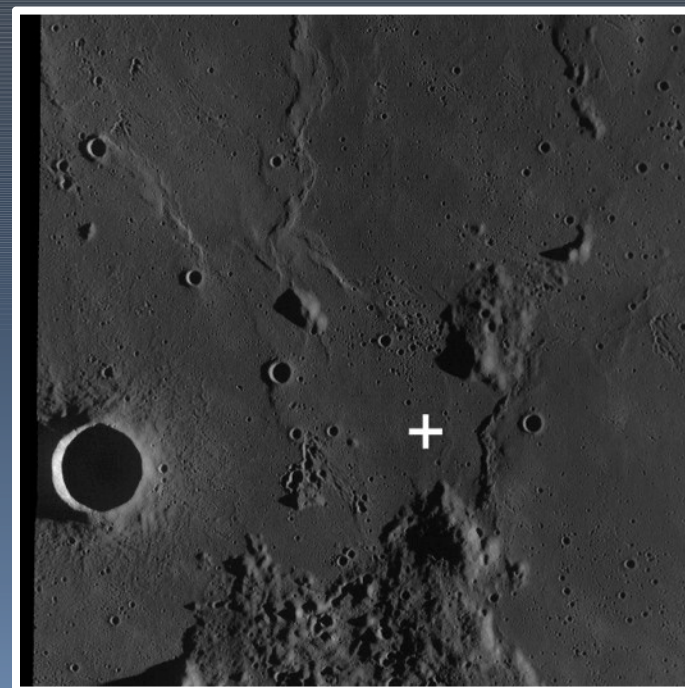
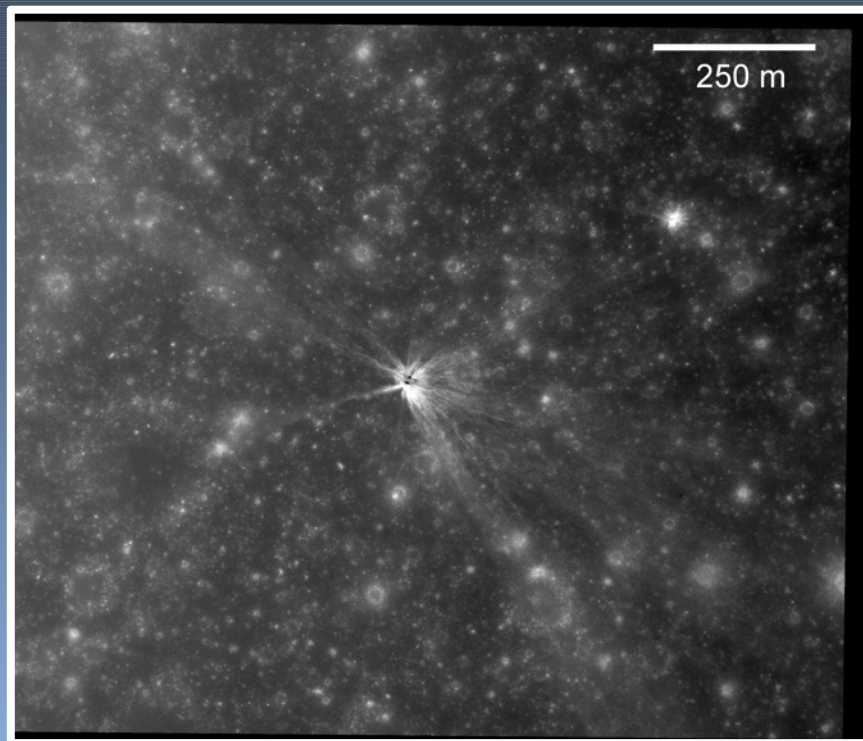
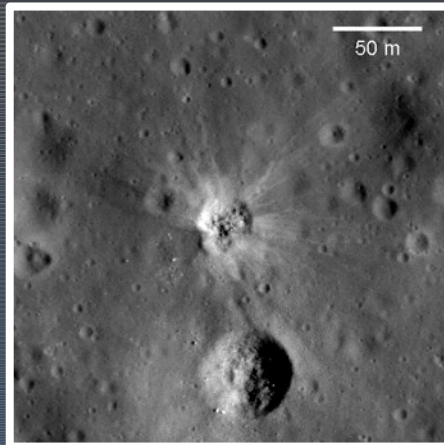
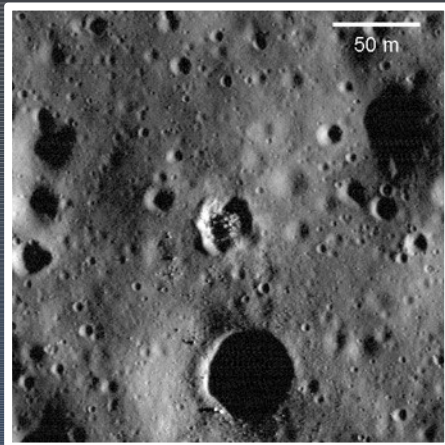
Fuel: LOX / LH<sub>2</sub>

Mass (Dry): 14,000 kg

Bulk Density: 20 kg m<sup>-3</sup> (0.02 g cm<sup>3</sup>)



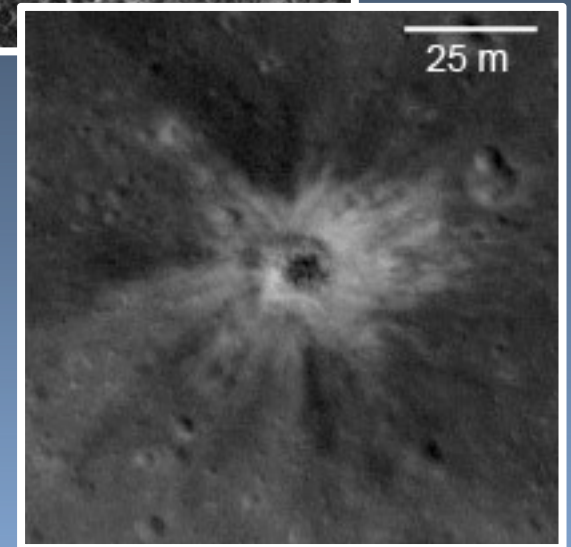
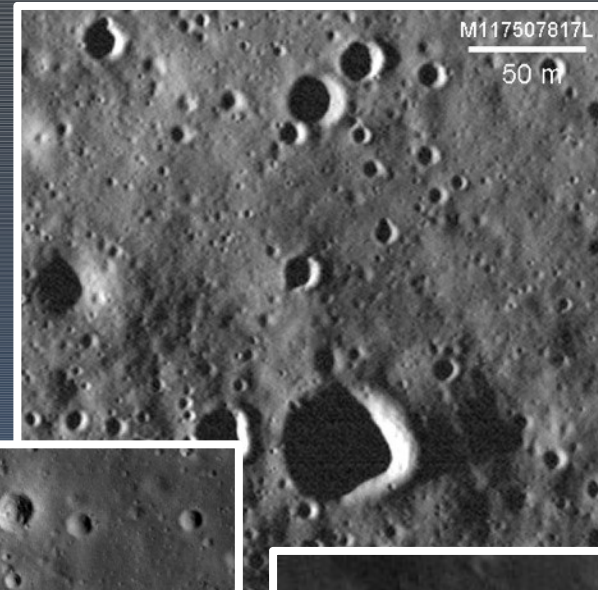
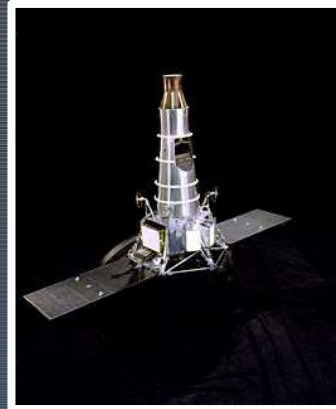
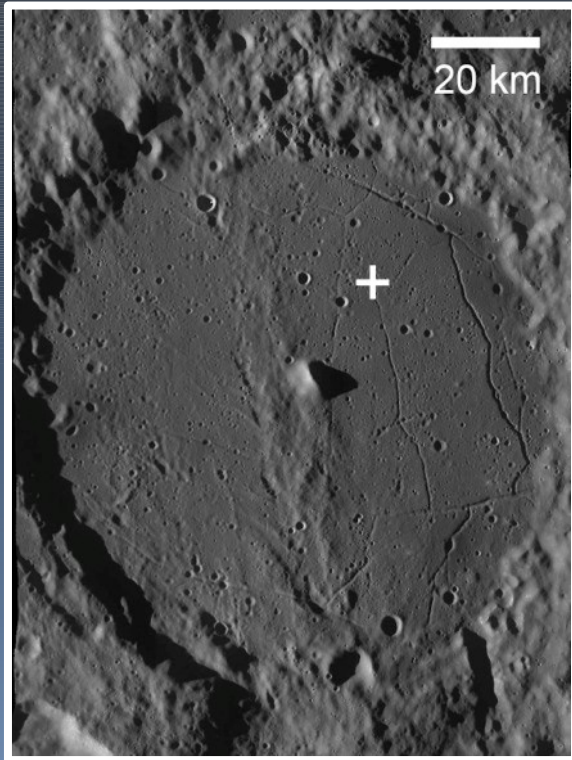
# Apollo 15 SIVB Crater



July 29, 1971  
Sinus Aestuum

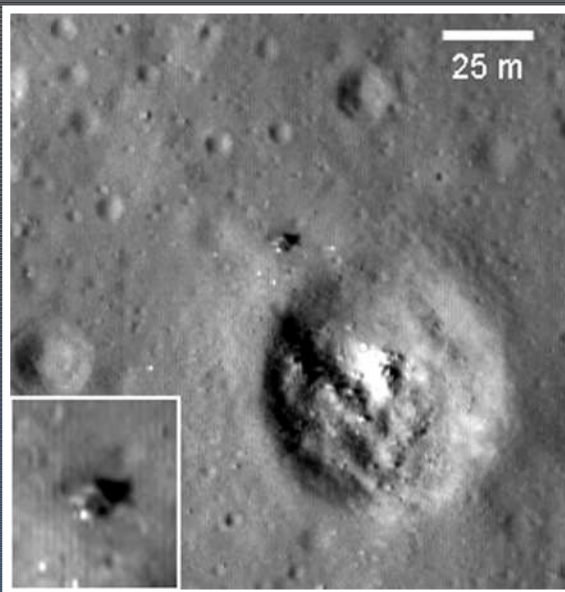


# Ranger 9 Impact Crater

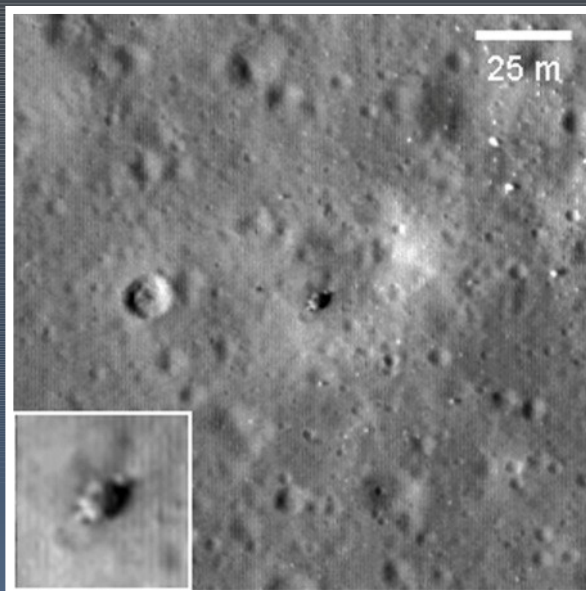


Alphonsus Crater Floor  
March 24, 1965

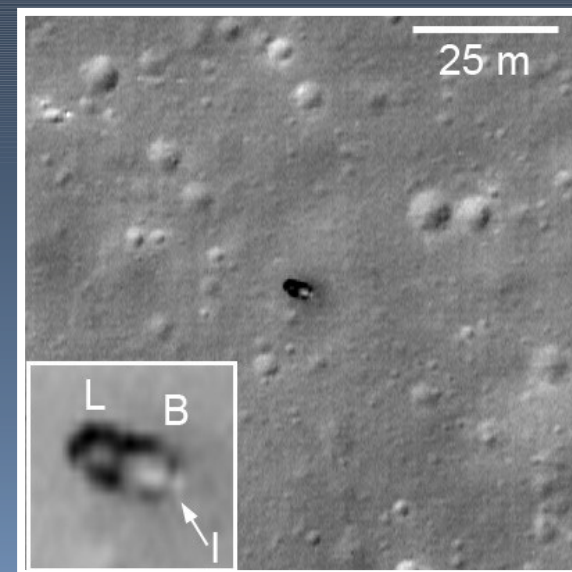
# Soviet Lunar Missions



Luna 24



Luna 23



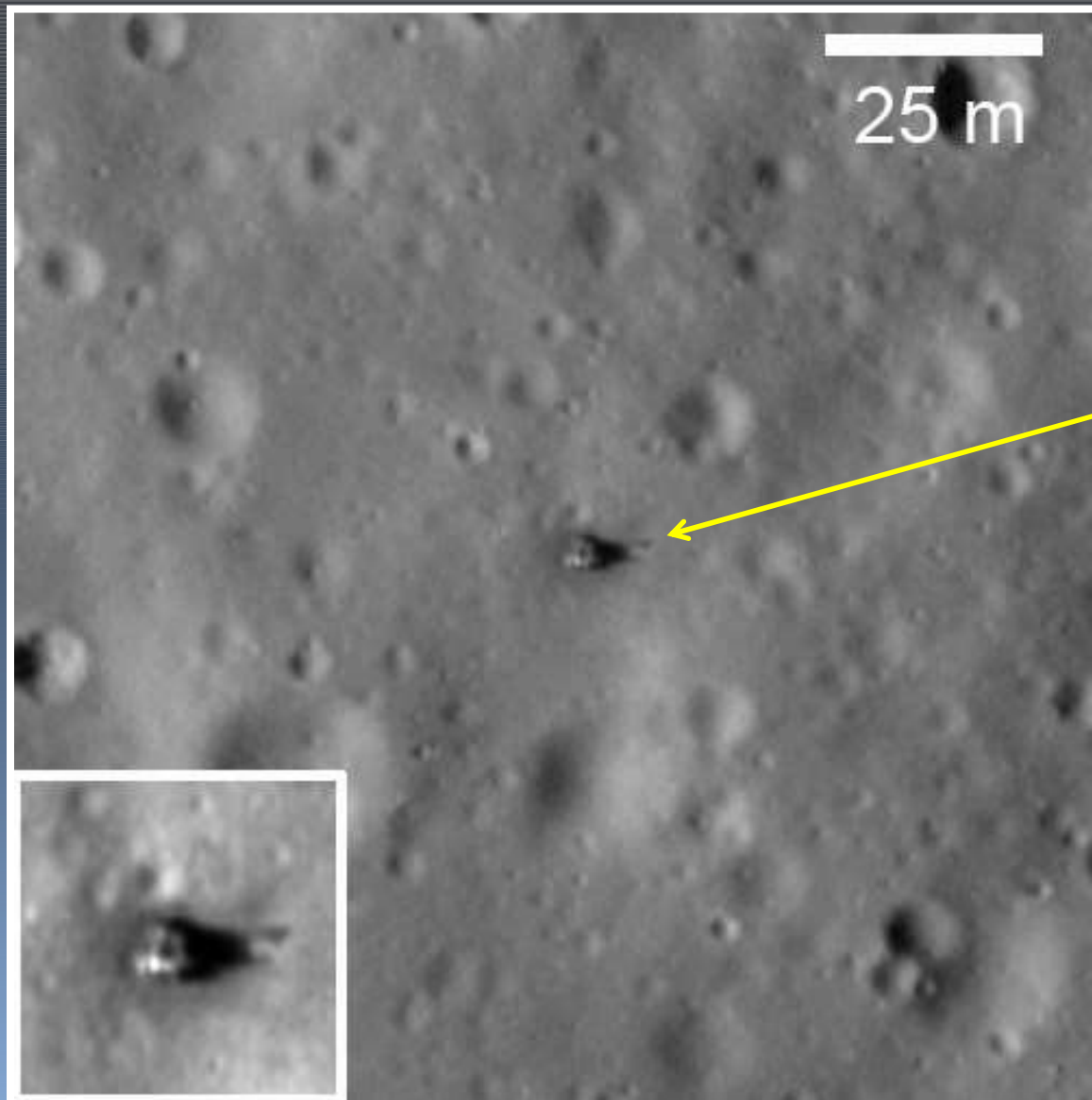
Lunokhod 2



Luna 23

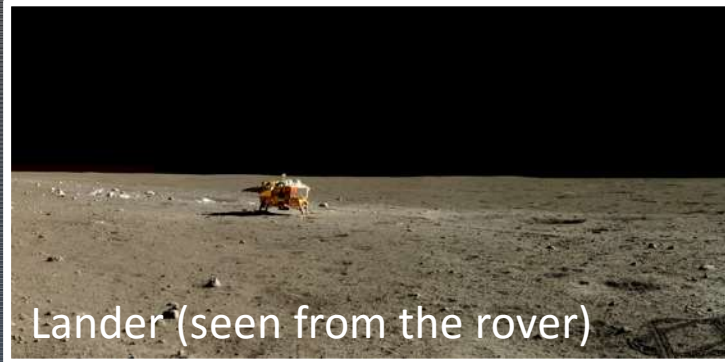


# Luna 20

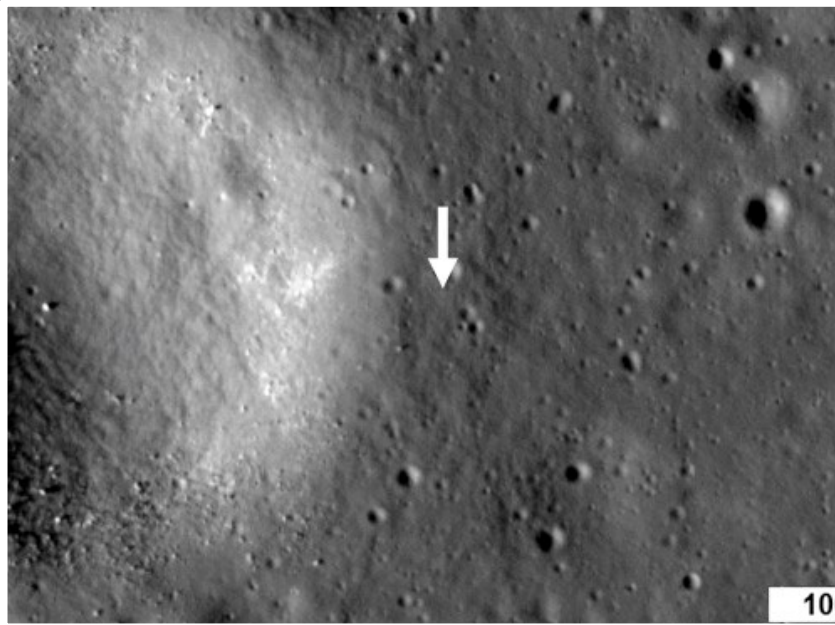


You can see the shadow of the sampling mechanism

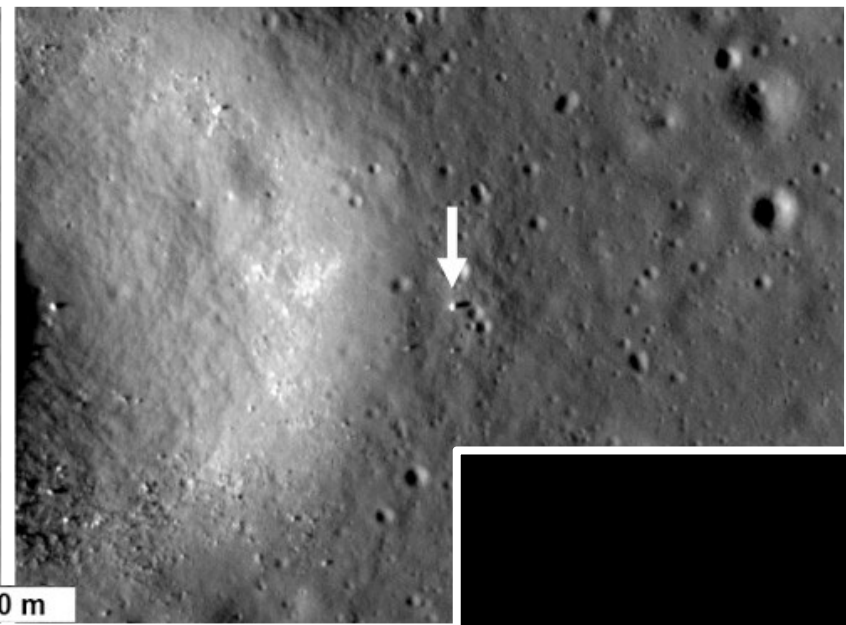
# Chang'E



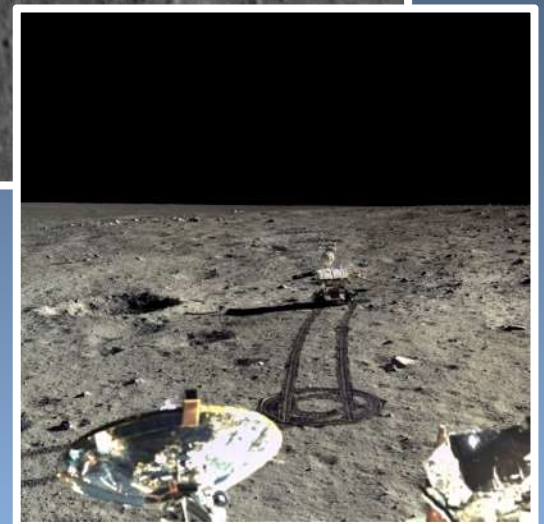
Lander (seen from the rover)



Before landing



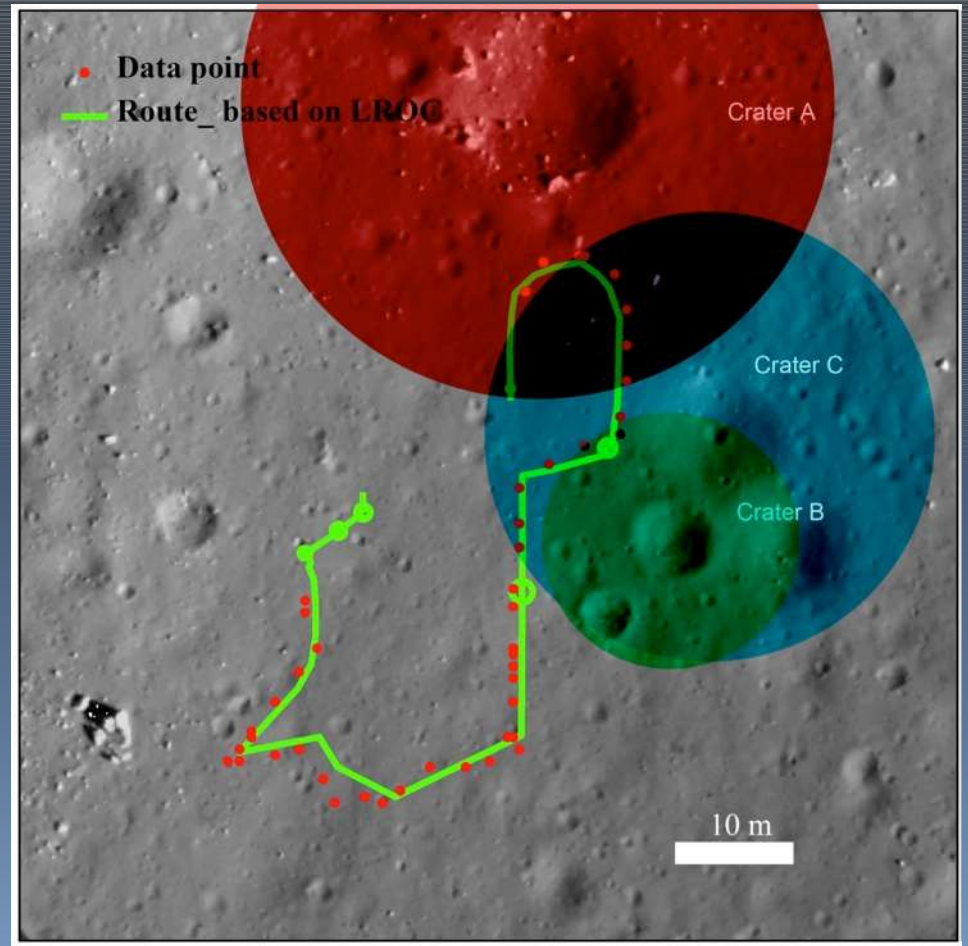
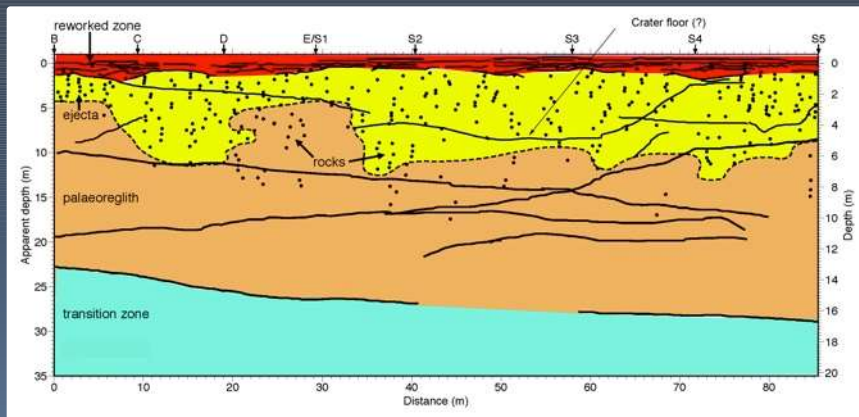
After landing



Rover seen from the lander

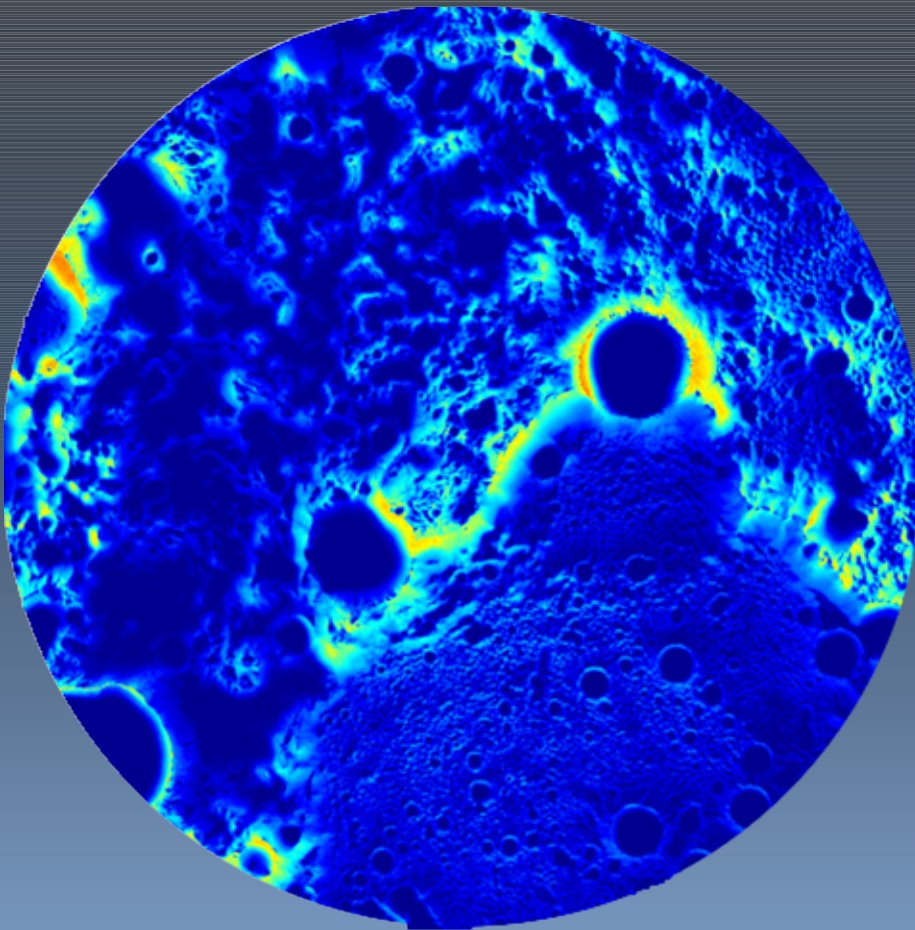


# Yutu - Ground Penetrating Radar

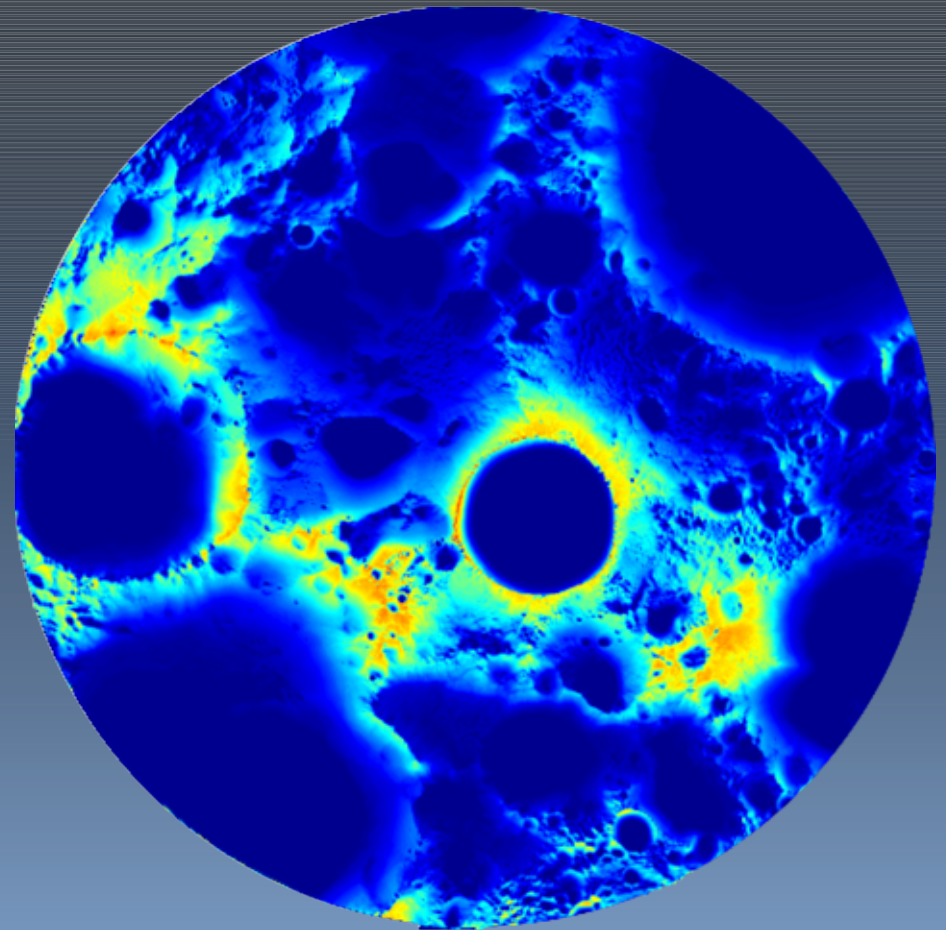


# WAC Illumination Maps

Feb 2010 - Jan 2011

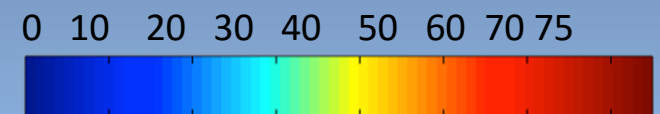


North Pole



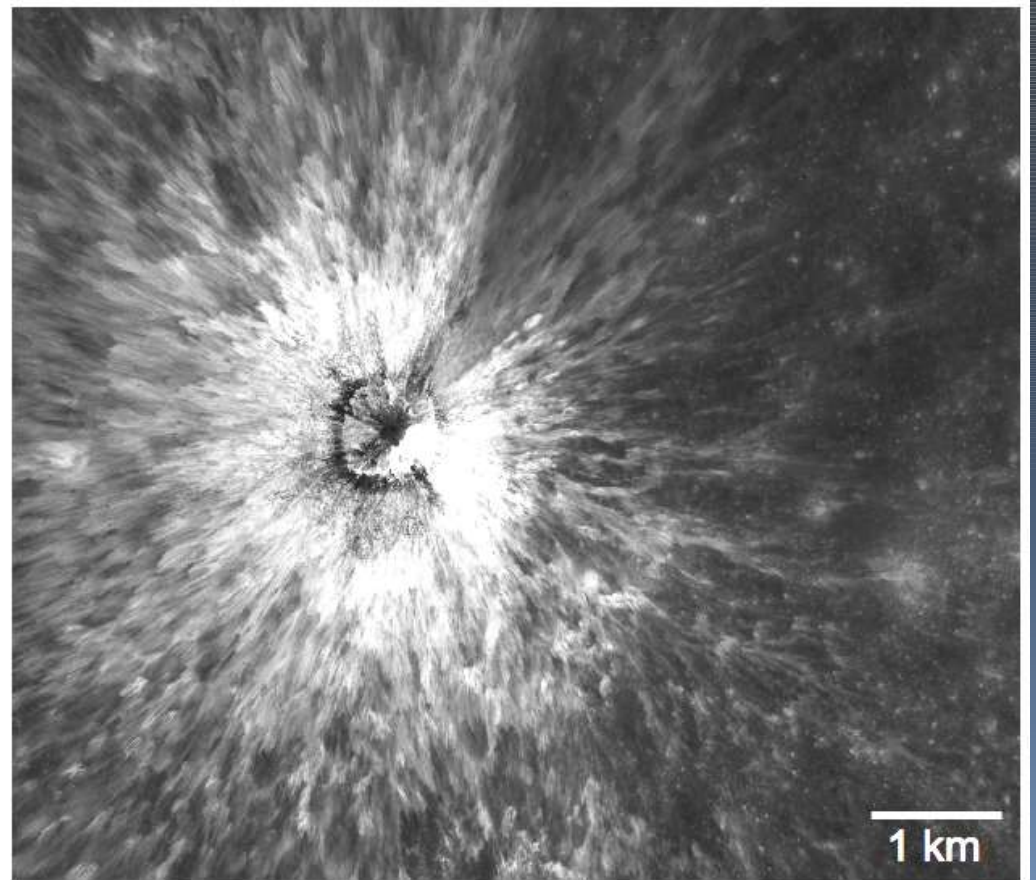
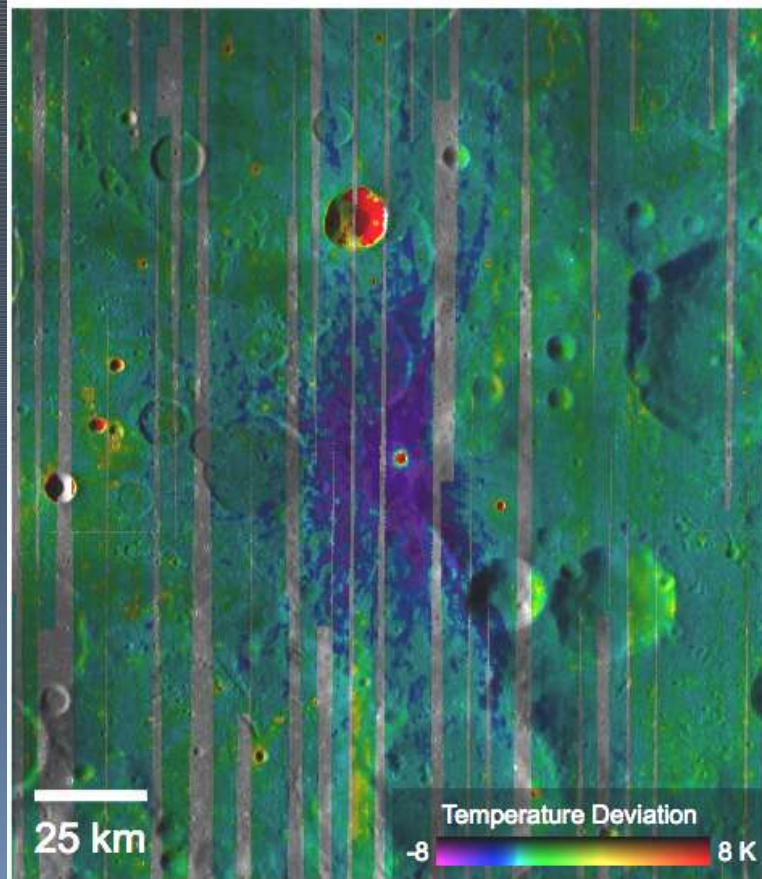
South Pole

The illumination maps extend from 88°N/S to the pole.



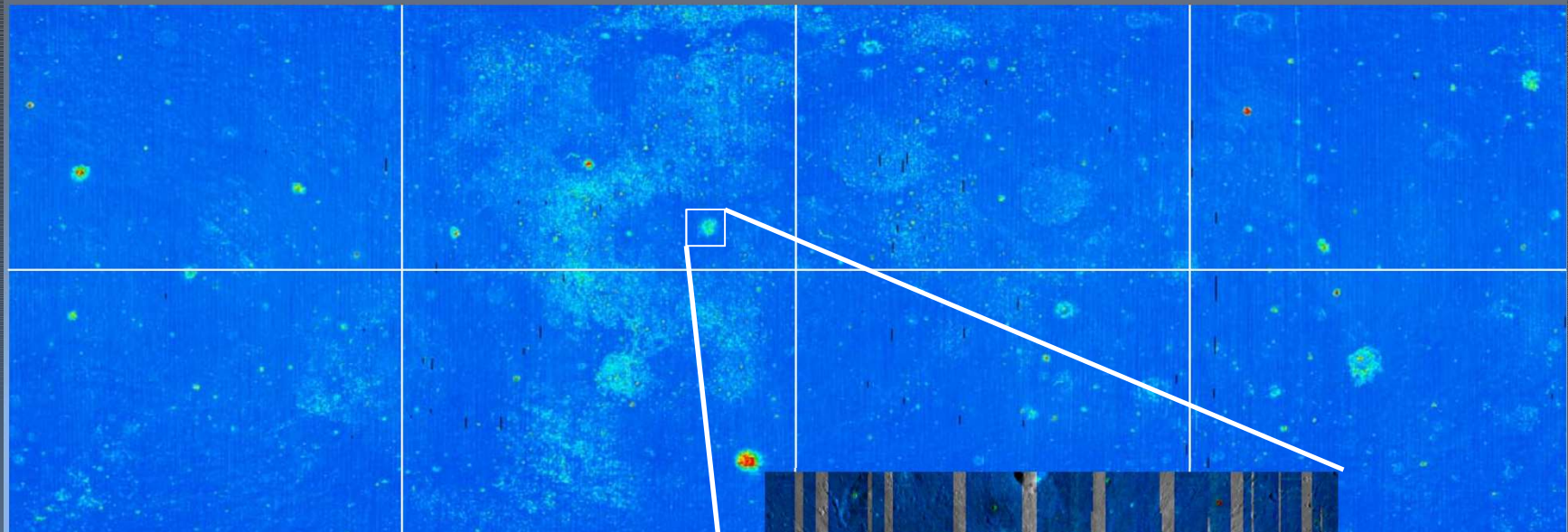


# Diviner Thermal Anomalies

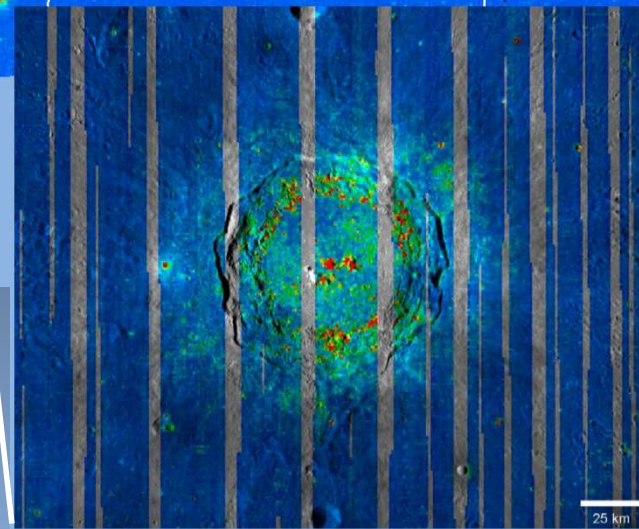


Large (100's of crater radii) regions around some fresh craters are unusually cold at night.

# Surface Block Abundance



Rock Concentration (0-0.05)

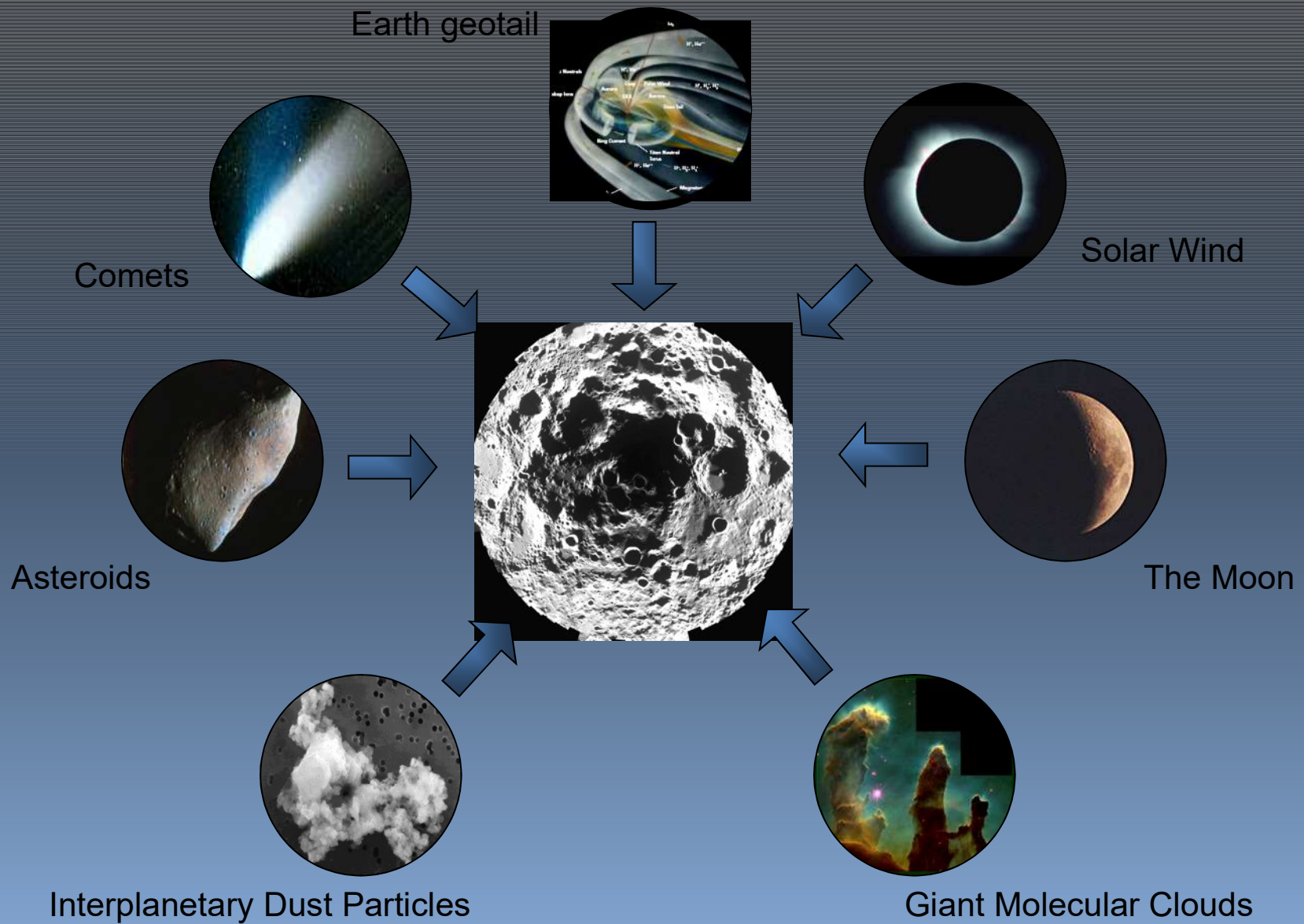


Copernicus crater

Rock Concentration 0 0.05

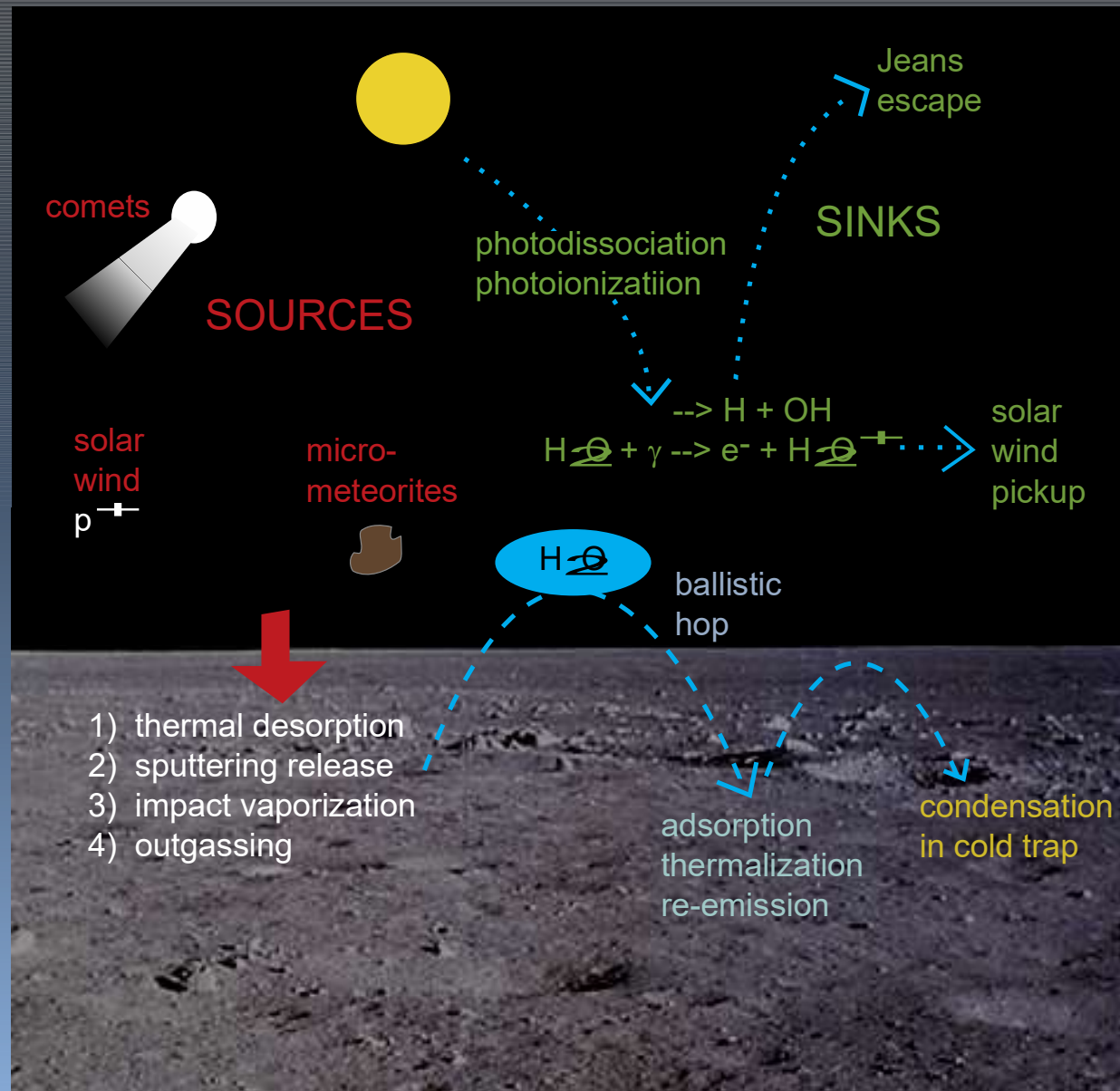


# Sources of Lunar Polar Volatiles



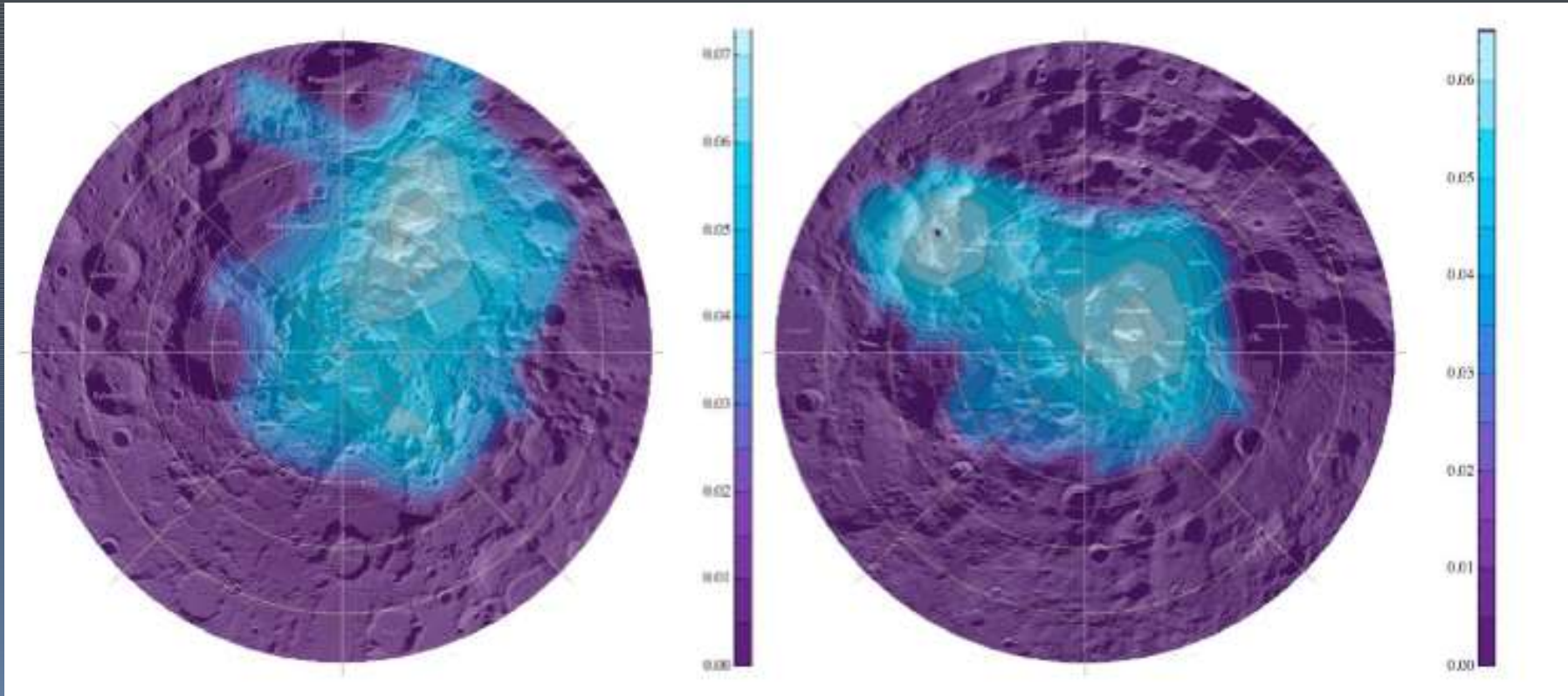
Modified from Lucey (2001)

# Sources and Migration of Volatiles to Cold Traps





# Polar Volatiles - Hydrogen



Neutron data – Hydrogen abundance (H, OH, H<sub>2</sub>O, C<sub>2</sub>H<sub>6</sub>O)

Average 0.01% wt. water equivalent hydrogen poleward of 80° and in the top 1-2 m.

If water, it amounts to  $9.8 \times 10^{10}$  kg (1/1000 of Lake Tahoe).

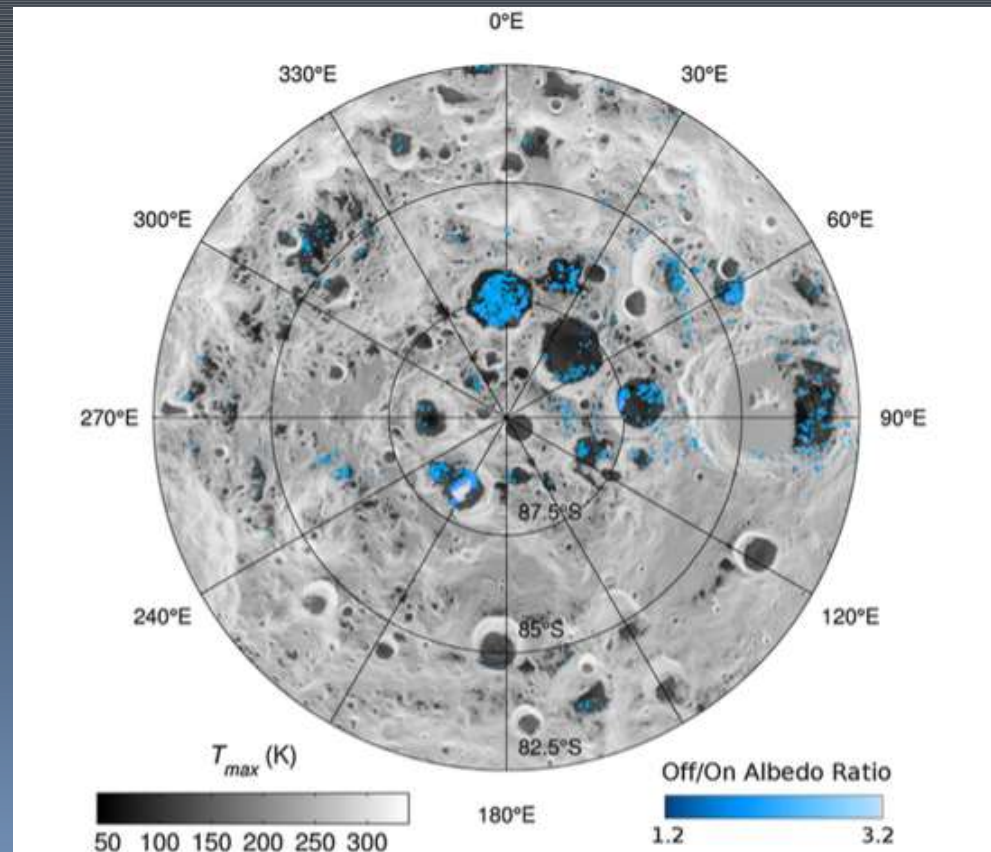
# Polar Volatiles – Water Frost

Heterogeneous lateral distribution

Surface frost

Signature of water ice at 165 nm - analysis shows surface frost is not evenly distributed in cold regions.

LRO LAMP data - low illumination with very coarse spectral binning, thus are better as supporting data than standing alone.



Ultraviolet reflectance and temperature.



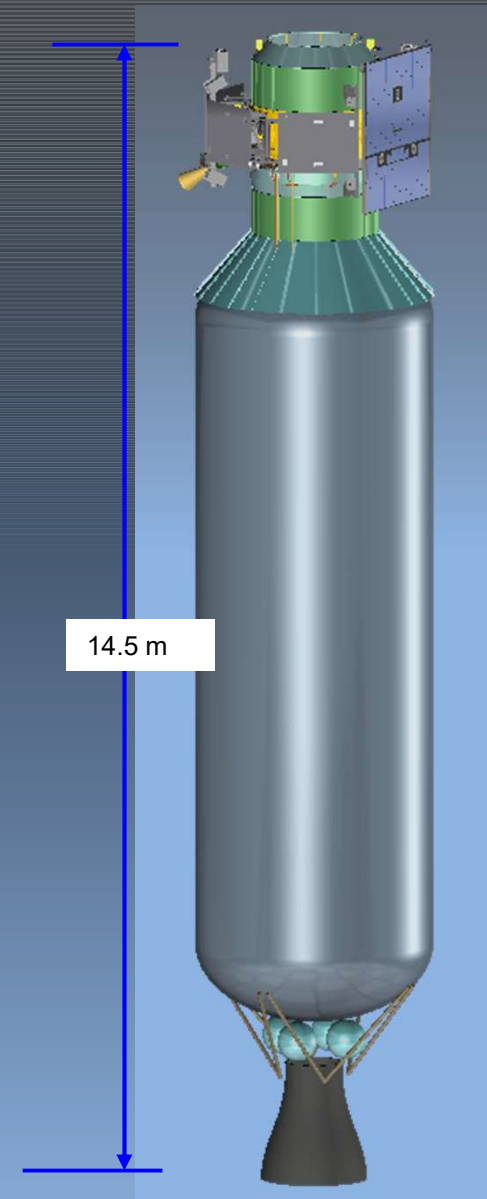
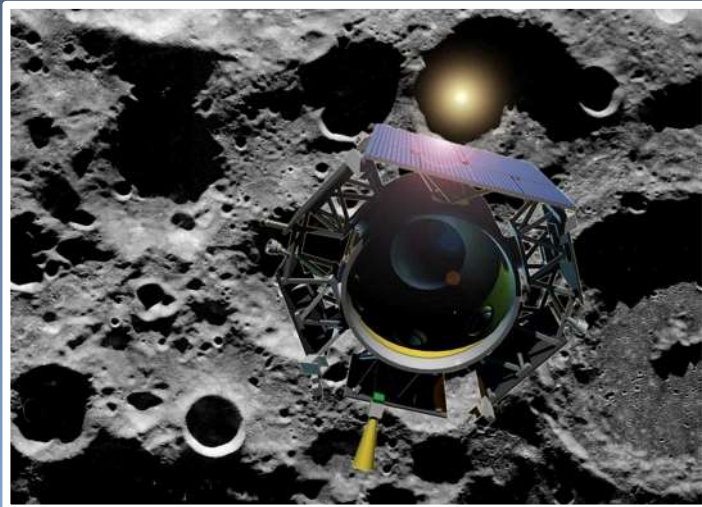
# Volatiles - LCROSS Mission

## Shepherding Spacecraft

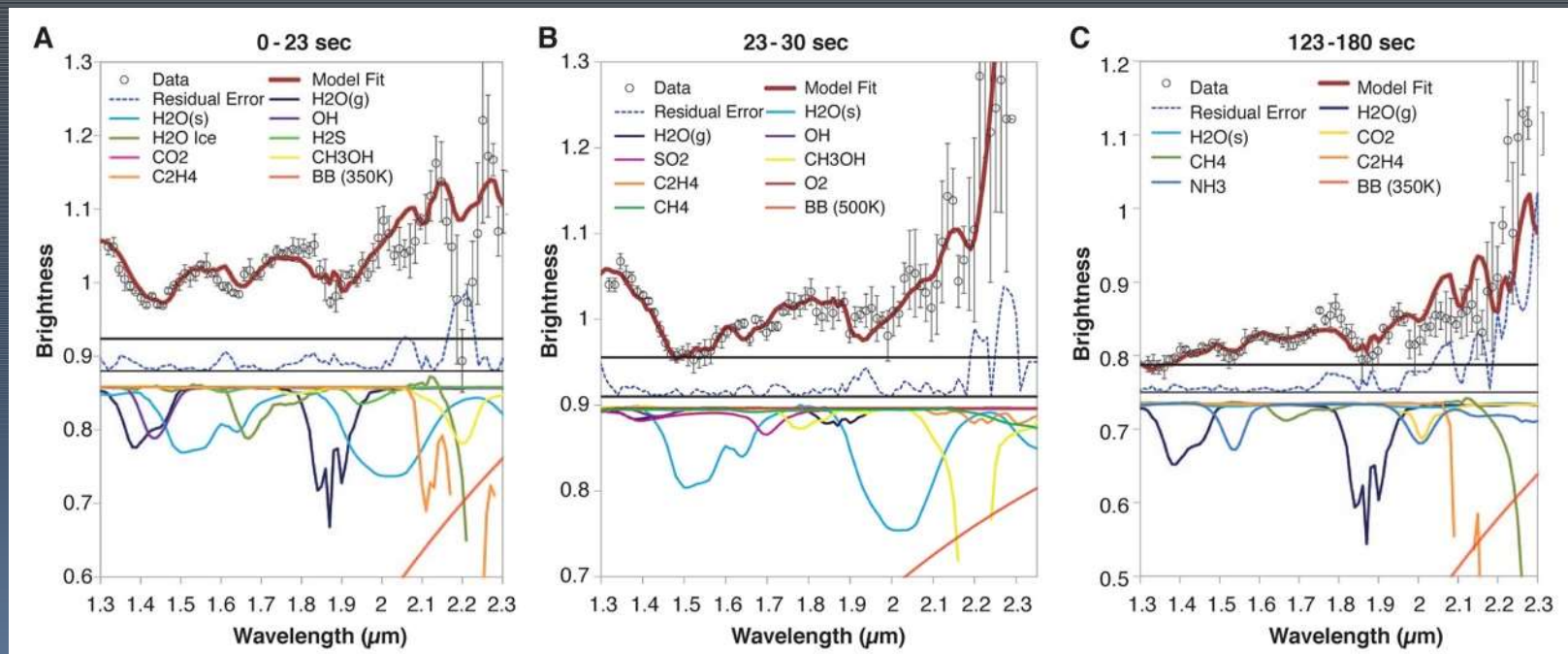
Guides and aims the Centaur to its target and carries all of the critical instrumentation.

## Centaur

Upper Stage: provides the thrust to get us from Earth orbit to the Moon and will then be used as an impactor



# Volatiles – LCROSS Mission



LCROSS impact into Cabeus

Spectral identification of H<sub>2</sub>O in both solid and gas phases

High resolution spectra provide strong evidence for the chemical composition.

Lofted material into sunlight providing a strong illumination source.

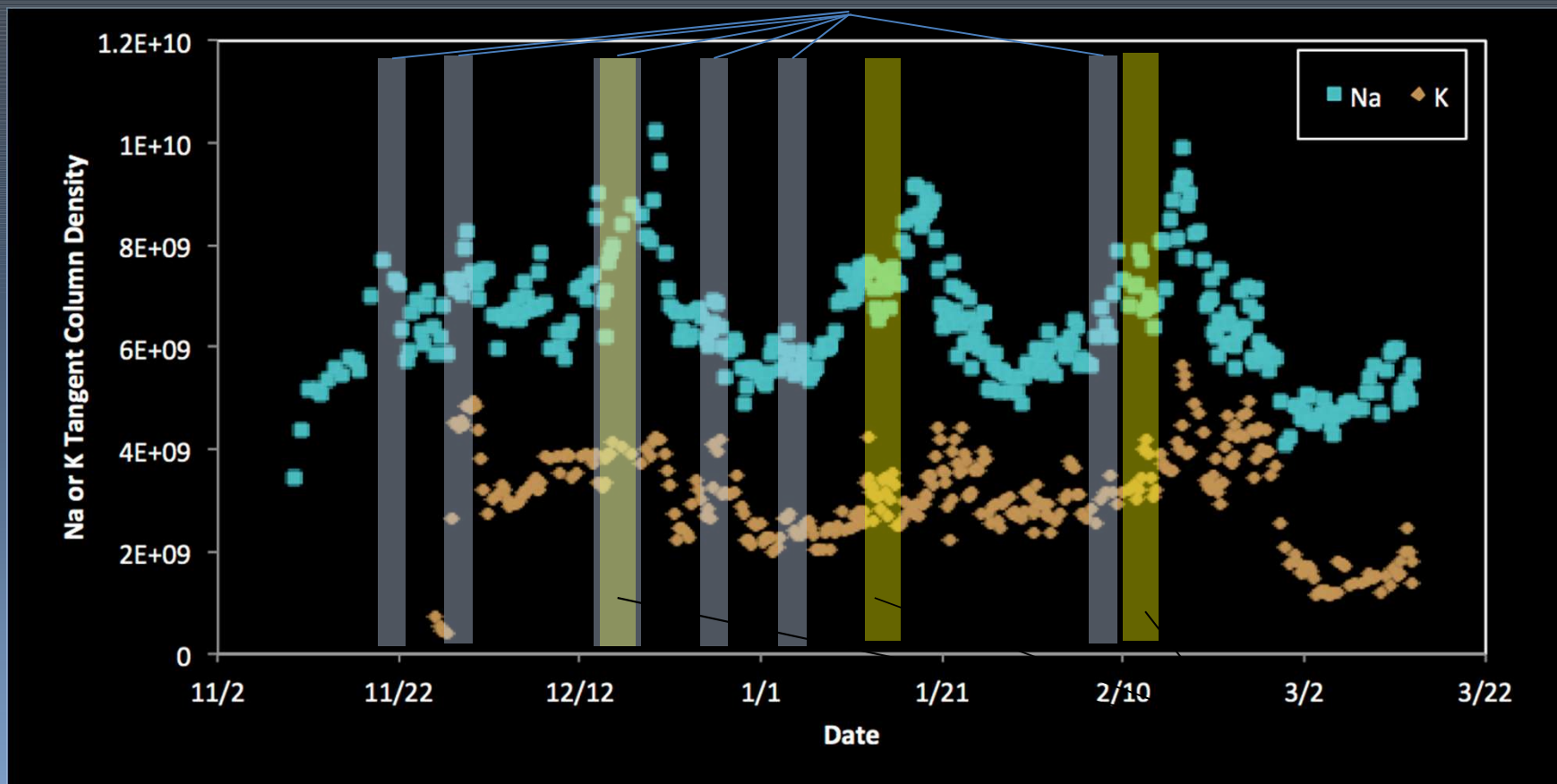
Question arises about any impact-induced chemistry and contributions from the impactor vs. the target material.



# Volatiles – LADEE / UVS

## Sodium and Potassium Observations

Meteoroid Showers

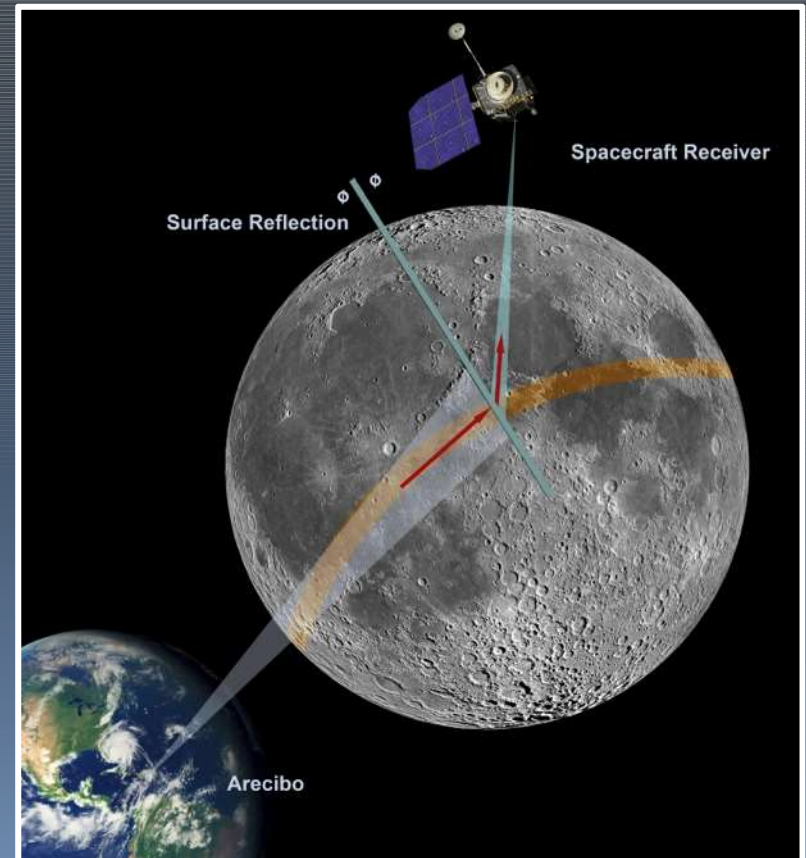


# Mini RF – Bistatic Radar Observations

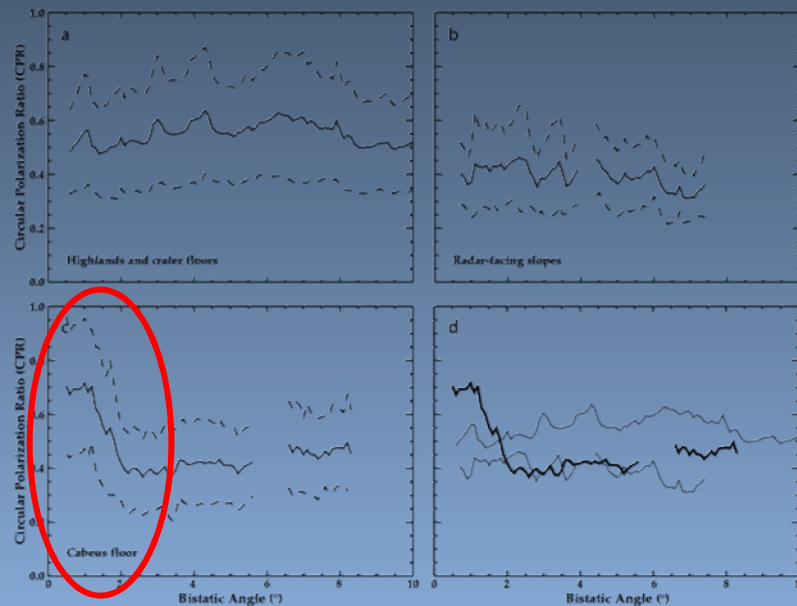
Mini-RF is currently operating in a bistatic mode, using the Arecibo Observatory or Goldstone as a transmitter

Experiment provides information on scattering properties of lunar materials as a function of beta (phase) angle

Within Cabeus Crater, there appears to be blocks of ice (~10 cm) in the subsurface.



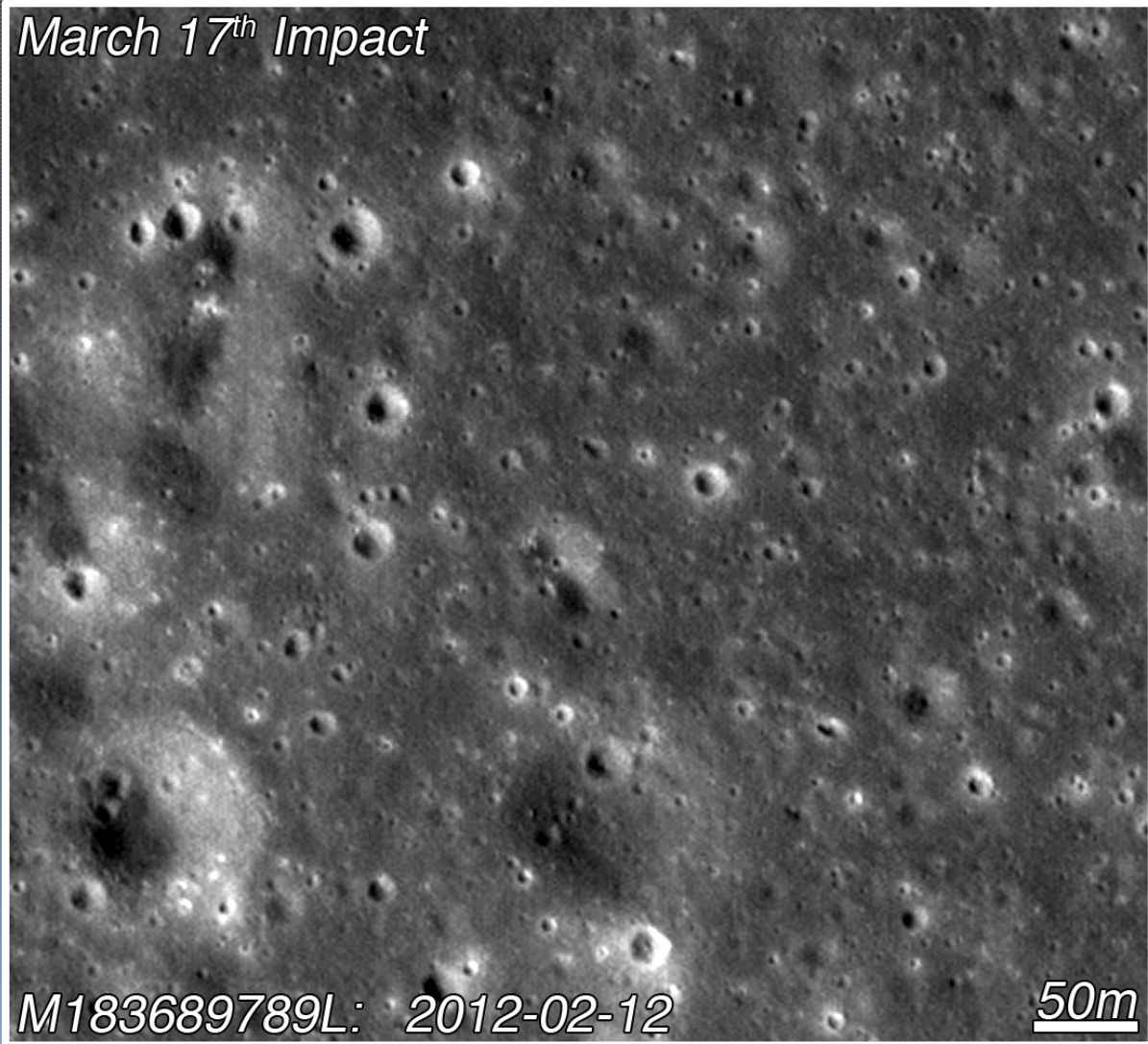
No ice



Ice

March 17, 2013

*March 17<sup>th</sup> Impact*



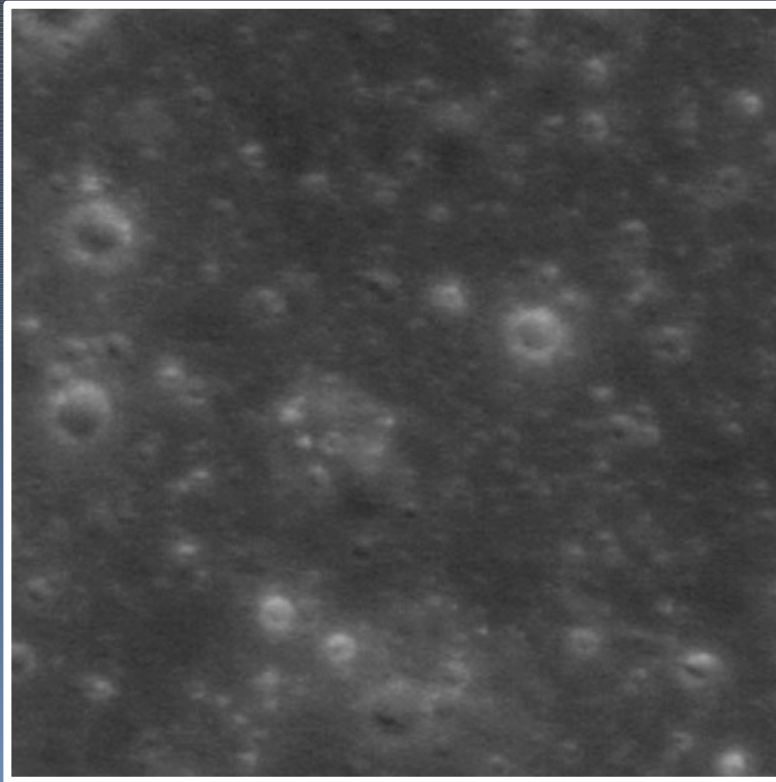
20 m diameter  
40 kg meteoroid  
0.3 - 0.4 m diameter  
5 tons of TNT

*M183689789L: 2012-02-12*

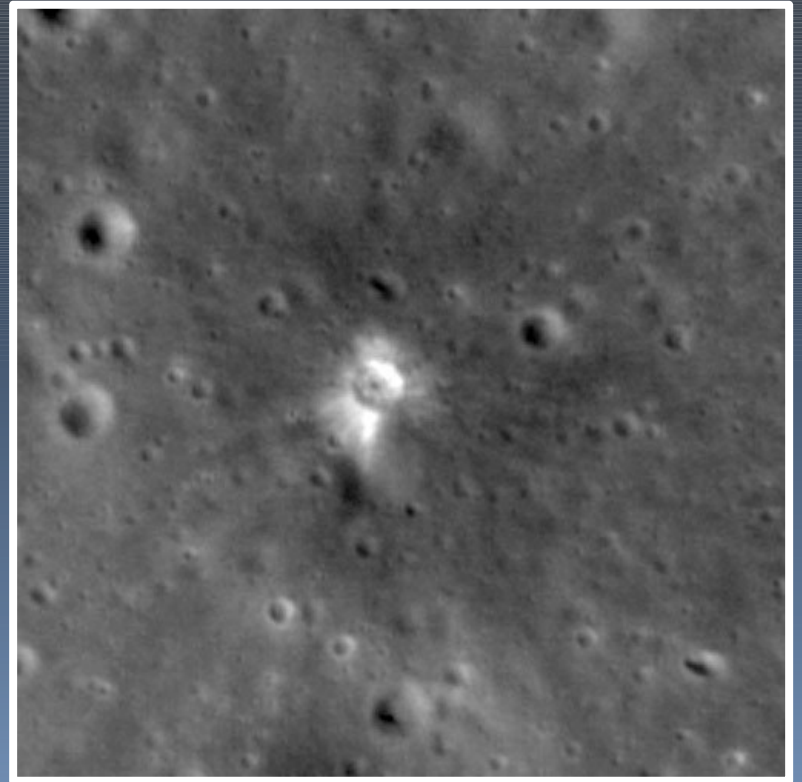
50m



March 17, 2013



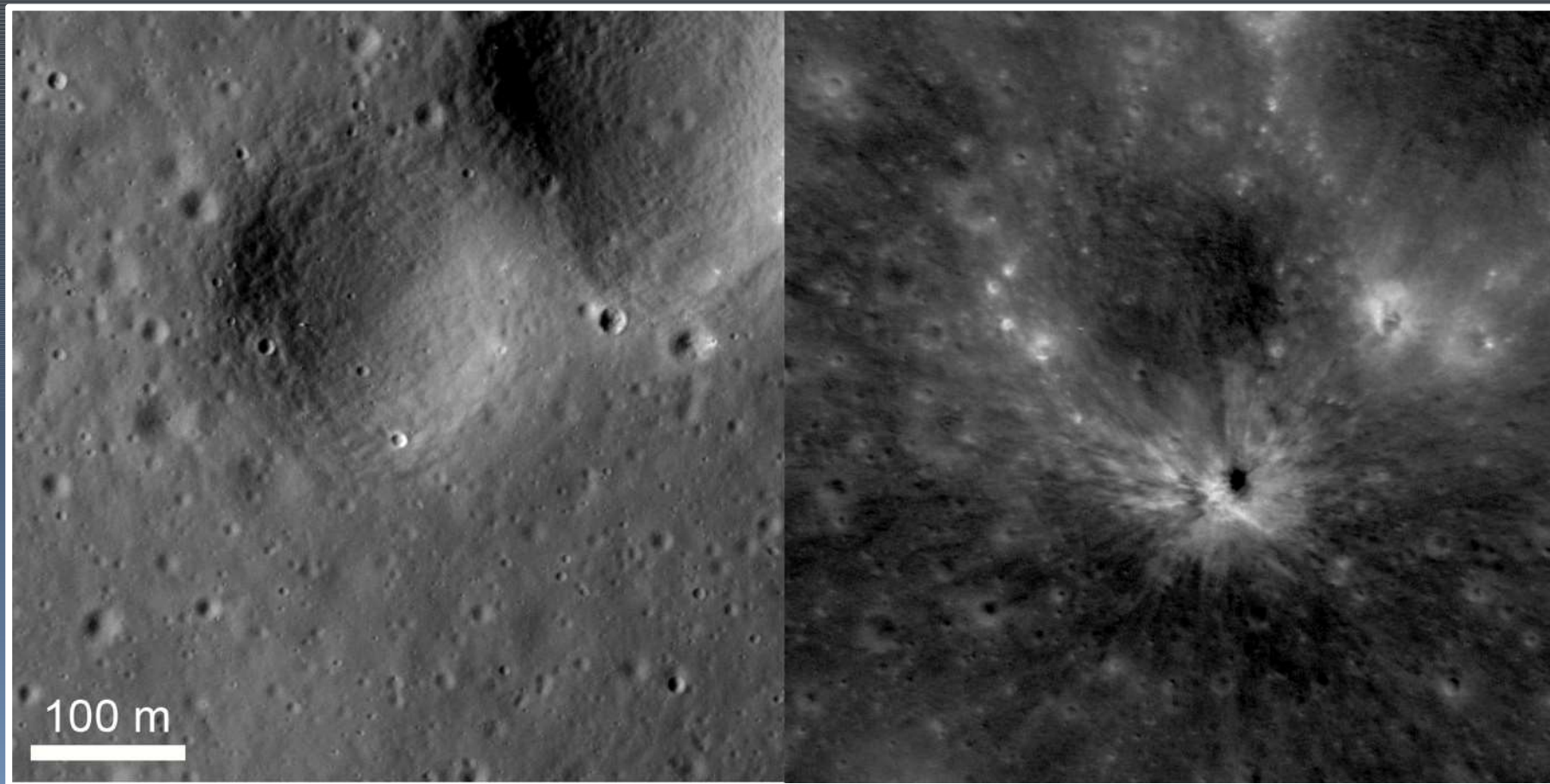
Before



After

20 m diameter  
40 kg meteoroid  
0.3 - 0.4 m diameter  
5 tons of TNT

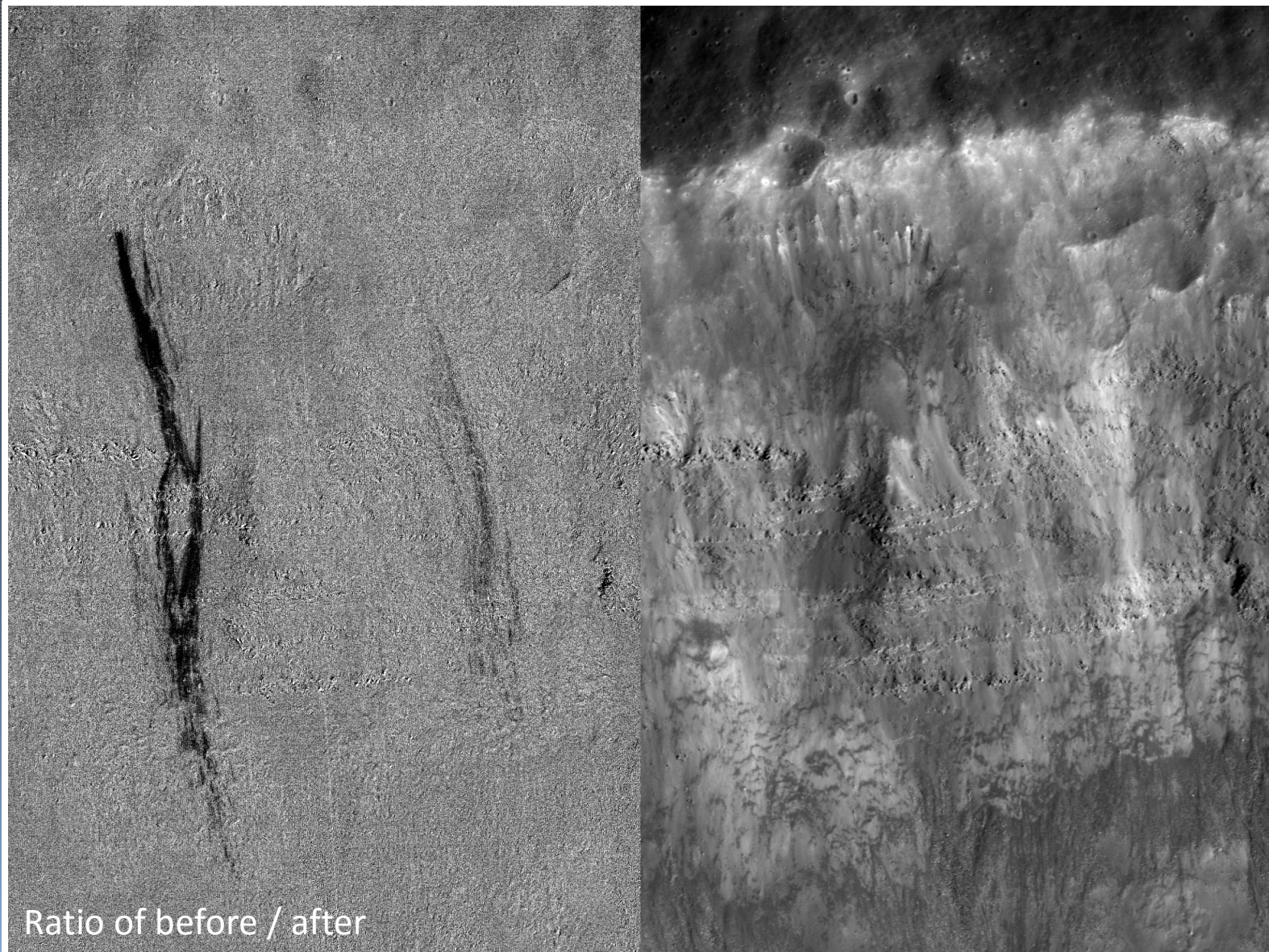
September 11, 2013



40 m diameter, 3-4 m deep, 0.6-1.4 m

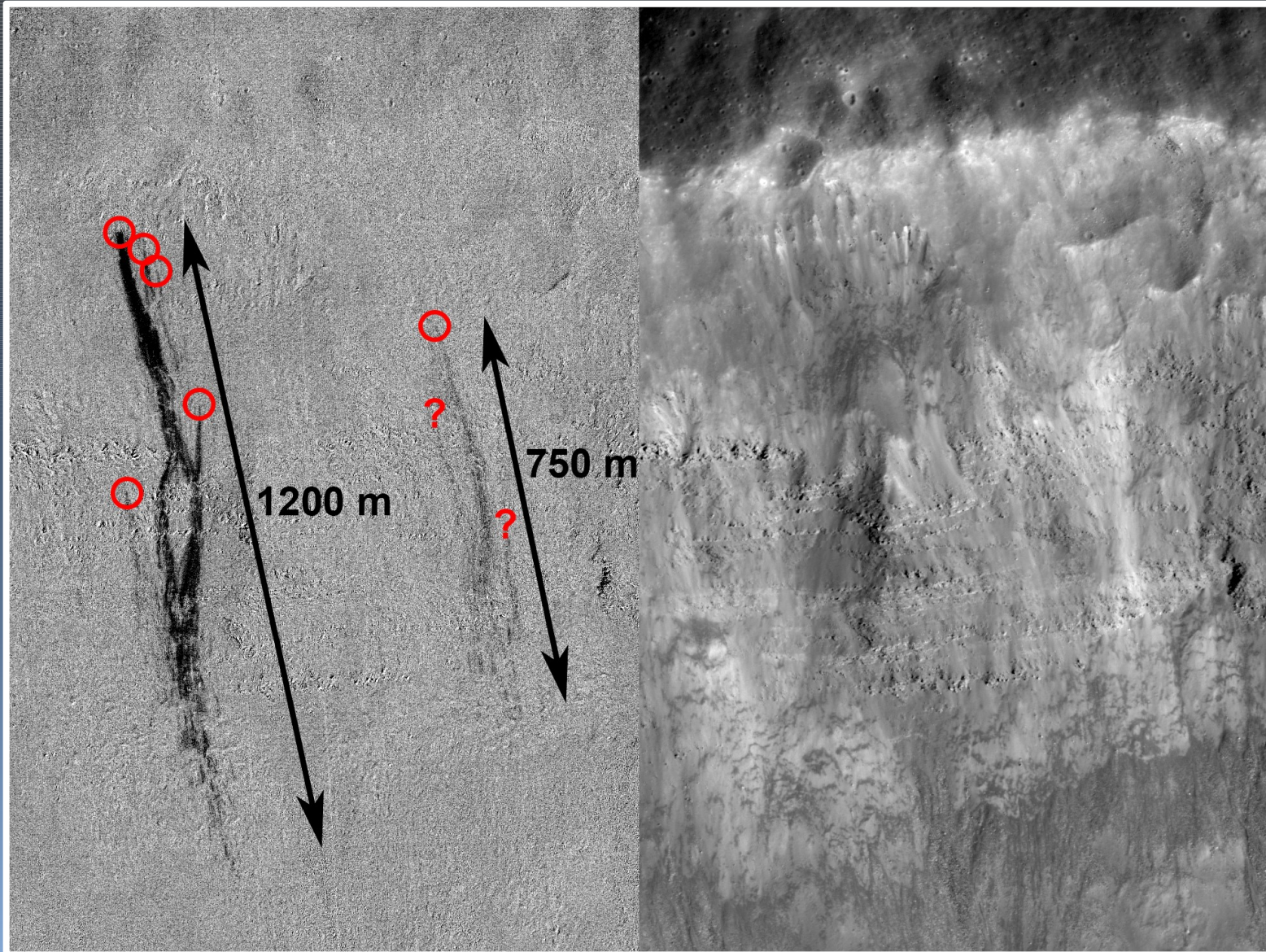


# Landslides



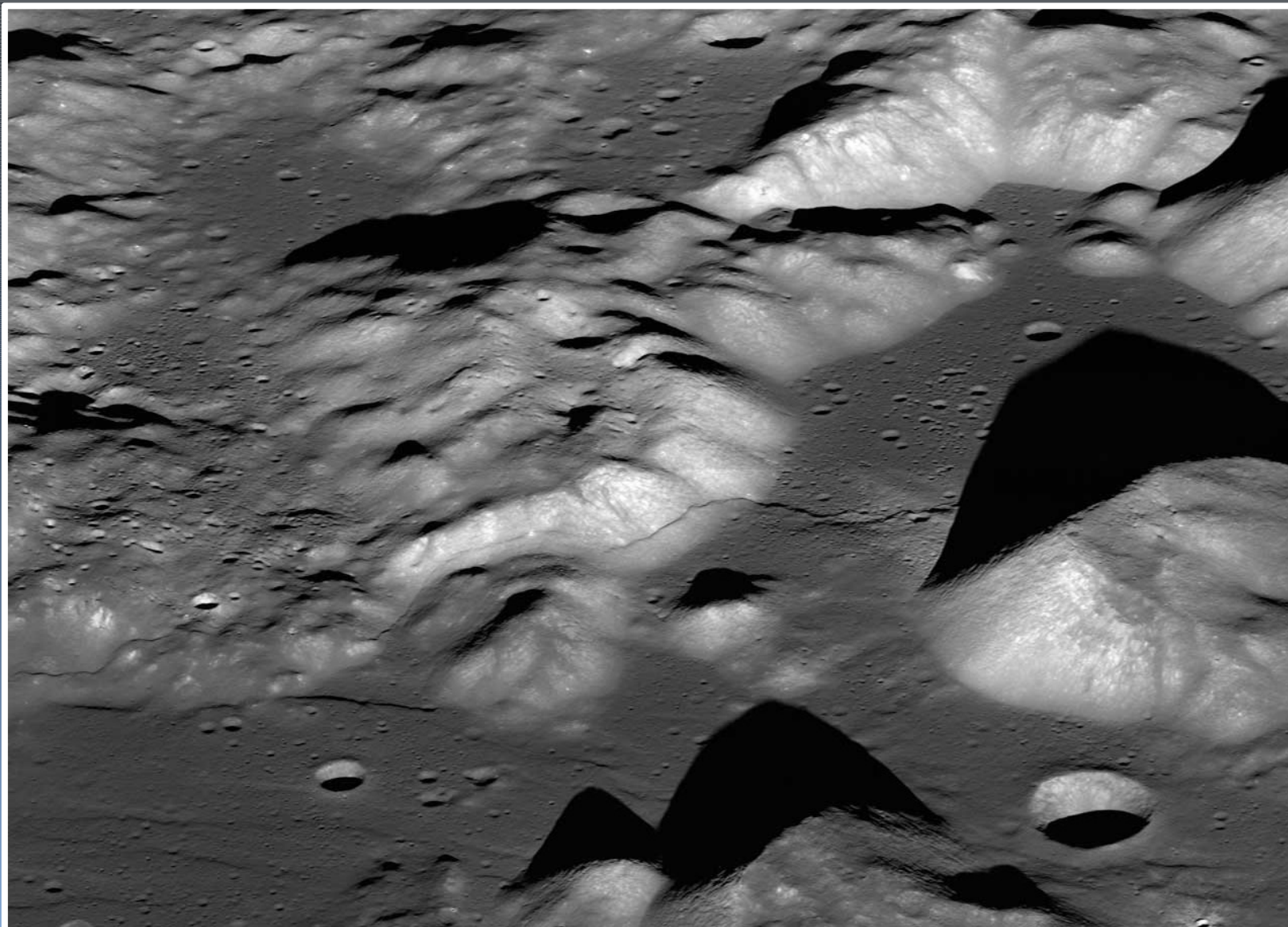


# Landslides





## Apollo 17 Site

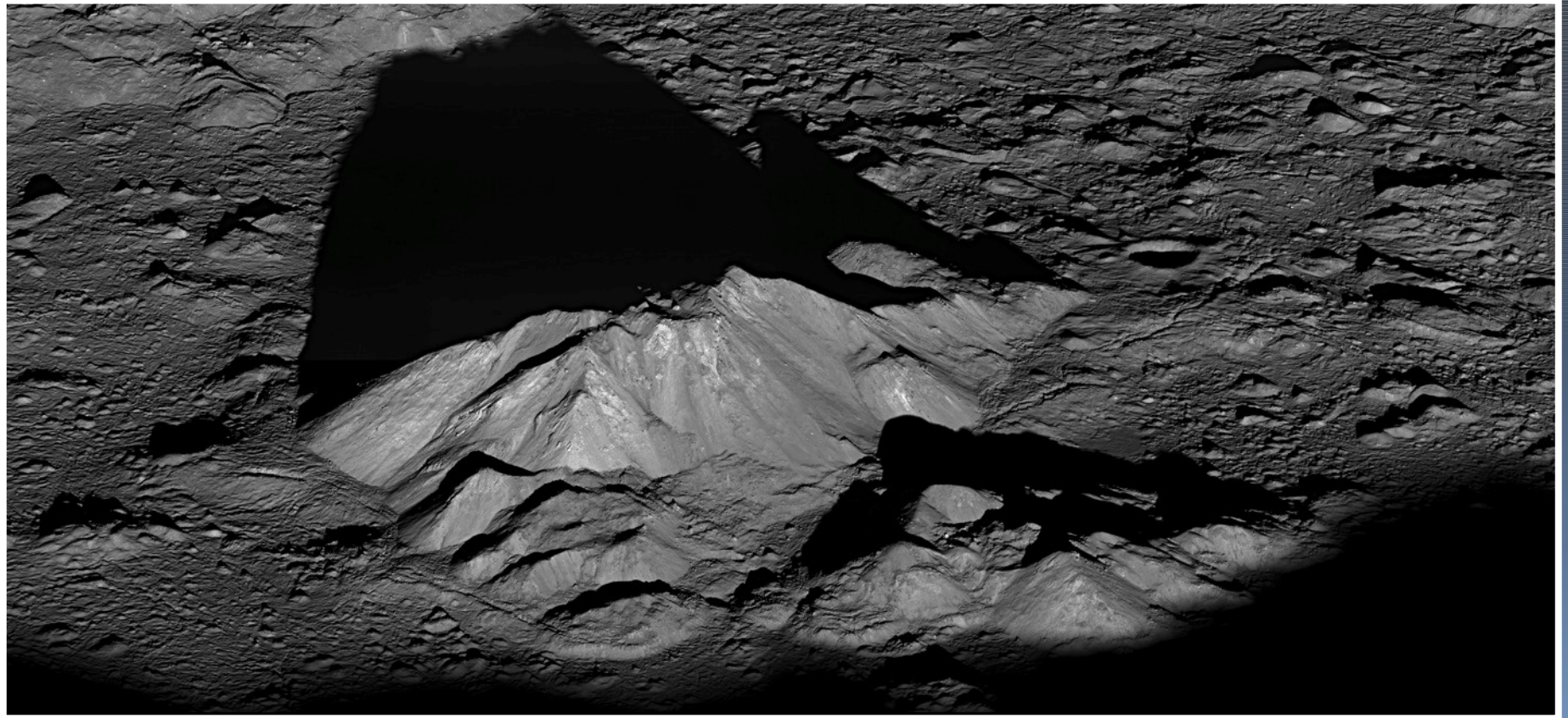


# Giordano Bruno

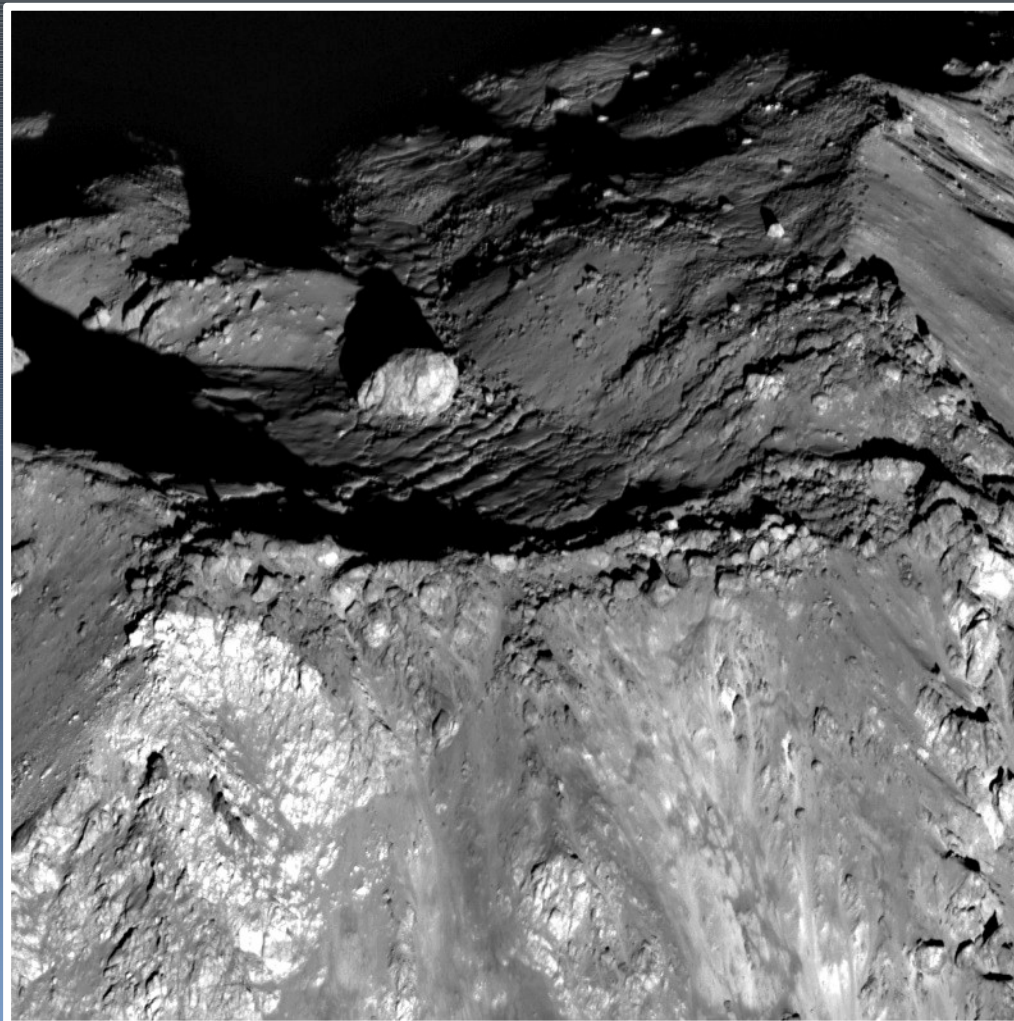




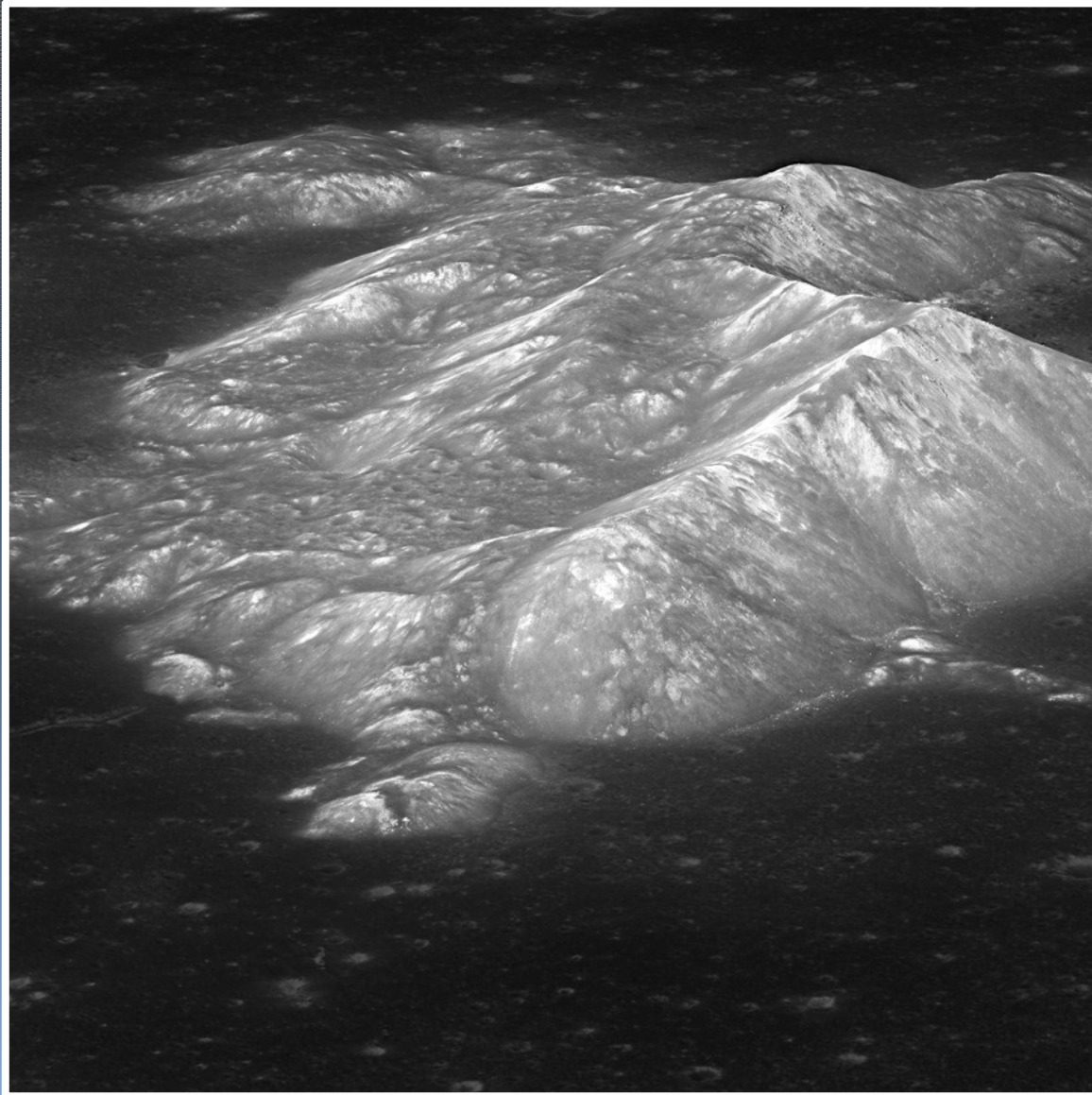
# Tycho



# Tycho

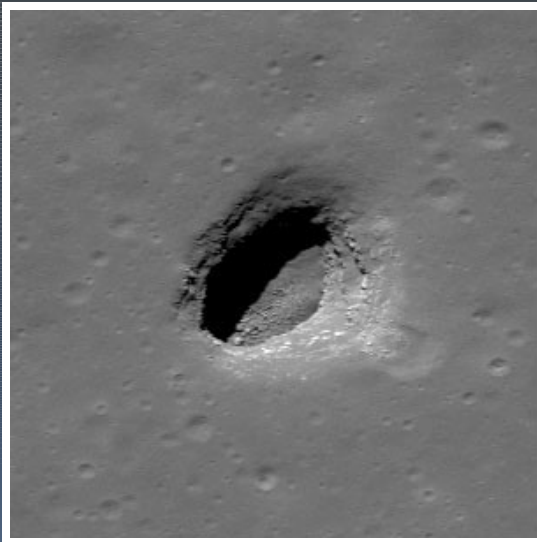


## Tsiolkovskiy Central Peak

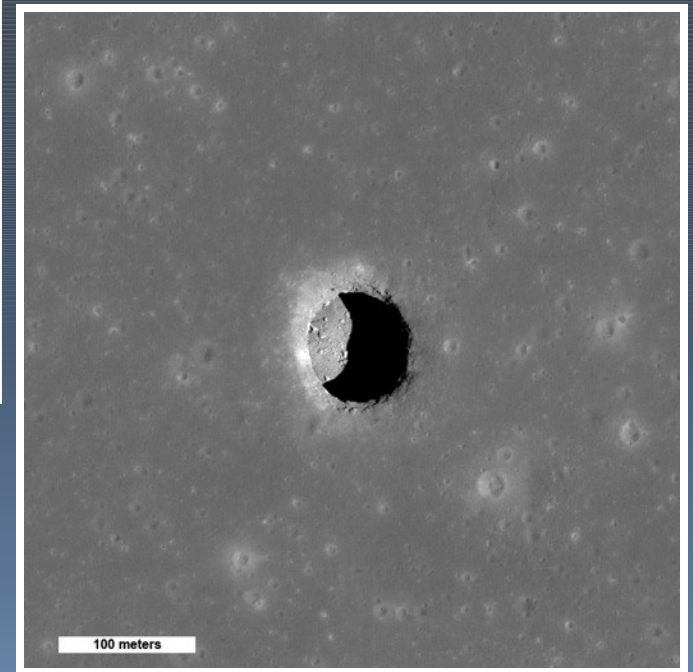
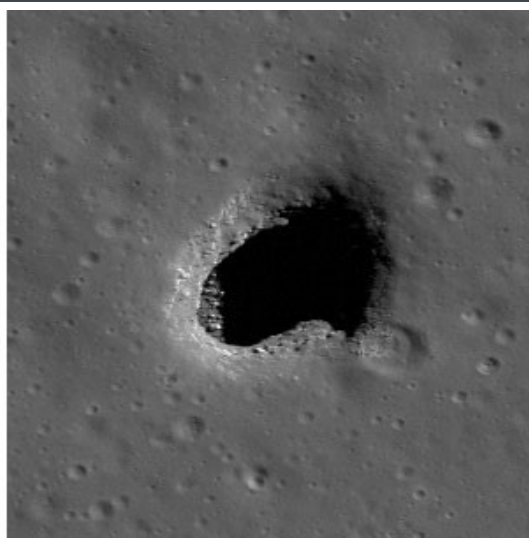




# Pits



Mare Ingenii



Mare Tranquillitatis

